

Ground-state properties of finite nuclei in relativistic Hartree-Bogoliubov theory with an improved quark mass density-dependent model*

Renli Xu (徐仁力)^{1†} Chen Wu (吴琛)^{2‡} Jian Liu (刘健)^{3§} Bin Hong (洪斌)^{4,5} Jie Peng (彭杰)⁶ Xiong Li (李雄)¹ Ruxian Zhu (朱如仙)¹ Zhizhen Zhao (赵至真)¹ Zhongzhou Ren (任中洲)^{4,5‡}

¹School of Artificial Intelligence and Information Technology, Jiangsu Province Engineering Research Center of TCM Intelligence Health Service, Nanjing University of Chinese Medicine, Nanjing 210023, China

²School of Physics and Electronic Information Engineering, Zhejiang Normal University, Jinhua 321004, China

³College of Science, China University of Petroleum (East China), Qingdao 266580, China

⁴School of Physics Science and Engineering, Tongji University, Shanghai 200092, China

⁵Key Laboratory of Advanced Micro-Structure Materials, Ministry of Education, Shanghai 200092, China

⁶School of Physics and Optoelectronic Engineering, Xiangtan University, Xiangtan 411105, China

Abstract: A relativistic Hartree-Bogoliubov (RHB) model based on quark-meson coupling is developed with a new parametrization derived from experimental observables. Using this model, we systematically investigate the ground-state properties of even-even nuclei spanning $8 \leq Z \leq 118$, including binding energies, quadrupole deformations, root-mean-square charge radii, two-nucleon separation energies, two-nucleon shell gaps, and α -decay energies. Comparisons with available experimental data confirm that this subnucleon-based RHB model reliably describes the ground-state properties of finite nuclei.

Keywords: relativistic Hartree-Bogoliubov theory, improved quark mass density-dependent model, deformation, binding energy

DOI: 10.1088/1674-1137/ae0997

CSTR: 32044.14.ChinesePhysicsC.50014105

I. INTRODUCTION

Nuclear energy density functionals (EDFs) grounded in density functional theory and the mean-field approach offer a universally applicable and powerful framework for describing the properties of finite nuclei and infinite nuclear matter. Over the past several decades, a variety of EDFs have been proposed, and they can be classified into two categories: non-relativistic and relativistic approaches. Within the non-relativistic domain, the most successful EDFs are the Hartree-Fock models based on density-dependent forces such as the Skyrme force for zero-range interactions [1, 2] and Gogny force for finite-range interactions [3]. As a relativistic version of the EDFs, the relativistic mean field (RMF) model describes nucleon interactions through the exchange of various virtual mesons [4–15]. Naturally incorporating relativistic effects, the RMF model treats spin-orbit coupling in a fully self-consistent manner without requiring additional adjustable parameters. By fitting a few free parameters to

experimental data, both non-relativistic and relativistic mean-field models enable an insightful understanding of a wide range of nuclear properties [16–47], including phenomena such as the nuclear skin and halo [16–27], shell effects [28–31], shape coexistence [32–37], exotic nuclear structures [38, 39], and proton and neutron drip lines [40–43]. These models also highlight the various properties of neutron stars [48–52].

Thus far, the RMF framework has evolved into a wide range of theoretical formulations. In the original Walecka model [4], nuclear interaction is described via the RMF theory with nucleons interacting through the exchange of scalar (σ) and vector (ω) mesons. Subsequently, Boguta and Bodmer extended the Walecka model by incorporating cubic and quartic self-interactions of the σ -meson, enhancing the description of the model of the nuclear matter incompressibility and surface characteristics [53]. The self-coupling and cross-coupling terms of the vector fields have also been systematically investigated [8, 54]. The introduction of isoscal-

Received 26 April 2025; Accepted 18 September 2025; Accepted manuscript online 19 September 2025

* Supported by the National Natural Science Foundation of China (12535009, 11605093, 12475135), and by the Natural Science Youth Fund of Jiangsu Province (BK20161036)

[†] E-mail: renlixu@njucm.edu.cn

[‡] E-mail: wucheno@fd@gmail.com

[§] E-mail: liujian@upc.edu.cn

[‡] E-mail: zren@tongji.edu.cn

©2026 Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd. All rights, including for text and data mining, AI training, and similar technologies, are reserved.

ar-isovector (ω - ρ) coupling significantly affects the density dependence of the symmetry energy [13, 55–57]. Although the standard approach outlined above is widely used, several studies have extended it to explore modifications and alternatives. For example, the derivative couplings of the scalar and vector fields within the standard RMF framework have been proposed [58, 59]. Relative to nonlinear meson self-interactions, relativistic models that explicitly incorporate density-dependent meson-nucleon couplings have seen significant advancements [60–64]. Besides these, the point-coupling model, which describes nucleon interactions via relativistic zero-range effective interactions, also marks a key progression within the RMF framework [65–68].

Alternatively, various effective models incorporating sub-nucleonic degrees of freedom have been successfully developed. A prominent example is the quark-meson coupling (QMC) model, originally formulated by Guichon in 1988 [69]. This innovative approach couples meson fields directly to confined quarks within nucleons, thereby establishing a theoretical framework that enables the systematic investigation of nucleon modifications in the nuclear medium and provides critical insights into the fundamental properties of nuclear matter [70–74]. Other theoretical frameworks based on quark-meson degrees of freedom include the chiral $SU(3)$ quark model [75–78], quark mean field (QMF) model [79–81], and improved quark mass density-dependent (IQMDD) model [82–84], among others. The IQMDD model is developed on the basis of the quark mass density-dependent (QMDD) model proposed by Fowler, Raha, and Weiner [85]. According to the QMDD model, the masses of the up (u), down (d), and strange (s) quarks, as well as their corresponding antiquarks are given by

$$m_q = \frac{B}{3n_B} \quad (q = u, d, \bar{u}, \bar{d}), \quad (1)$$

$$m_{s,\bar{s}} = m_{s0} + \frac{B}{3n_B}, \quad (2)$$

where n_B represents the baryon number density, m_{s0} represents the current mass of the strange quark, and B is the bag constant. Perturbative QCD fails to provide a confinement solution for quarks, and therefore, a confinement potential proportional to r (or r^2) is introduced into the quark system to prevent quarks from reaching regions infinitely distant or excessively large [79, 80]. In the QMDD model, the confinement mechanism is manifested by requiring the mass of isolated quarks to approach infinity, rendering the vacuum incapable of supporting them [86, 87]. As the volume approaches infinity or density tends to zero, the quark mass diverges, which is precisely the behavior characterized by equations (1) and (2). This confinement mechanism is analogous to that

in the MIT bag model [88–90]. Building on the QMDD framework, the IQMDD model extends the original formalism by systematically incorporating σ , ω , and ρ meson fields. This generalization establishes a self-consistent coupling between quark degrees of freedom and mesonic fields throughout the nuclear medium, enabling rigorous nuclear many-body calculations within the mean-field approximation [91].

The construction and operation of new-generation radioactive ion beam (RIB) facilities have produced an increasing number of nuclei far from β -stability valley, presenting new frontiers in nuclear physics that expand our understanding of nuclear phenomena from stable nuclei to exotic ones [92–94]. Exotic nuclei near drip lines are weakly-bound systems, where the Fermi surface of protons (or neutrons) is close to the particle continuum. This leads to an increased scattering of Cooper pairs into the continuum states as a result of pairing correlations, and furthermore, the conventional BCS model with the monopole pairing force can provide only a poor approximation in this case [95]. For weakly bound nuclei, a unified and self-consistent treatment of the mean field and pairing correlations becomes essential. The development of the relativistic Hartree-Bogoliubov (RHB) model has provided an effective approach to address this issue [33, 96–99]. In the RHB model, a unified description of particle-hole (ph) and particle-particle (pp) correlations is given, and the continuum effects are effectively considered in the coordinate representation [100–105]. In this paper, RHB calculations incorporating quark-meson coupling are presented within the IQMDD framework, and they serve as a natural extension of the IQMDD model. With this new model, we systematically calculated the ground-state properties of even-even nuclei with proton numbers ranging from $Z=8$ to $Z=118$ where experimental data are available. This remainder of this paper is organized as follows. The main formulas of the IQMDD+RHB model are presented in Section II. The model parameters and numerical results are presented and discussed in Section III. Finally, a summary is given in Section IV.

II. THEORETICAL FRAMEWORK

The effective Lagrangian density of the IQMDD model can be expressed as

$$\begin{aligned} \mathcal{L} = & \bar{\varphi}_q \left\{ \gamma^\mu \left[i\partial_\mu - g_\omega^q \omega_\mu - \frac{g_\rho^q}{2} \tau^q \cdot \rho_\mu - \frac{e}{2} \left(\frac{1}{3} + \tau_3^q \right) A_\mu \right] \right. \\ & + \frac{f_\omega g_\omega^q}{2M_N} \sigma^{\mu\nu} \partial_\nu \omega_\mu - m_q + g_\sigma^q \sigma \left. \right\} \varphi_q + \frac{1}{2} \partial^\mu \sigma \partial_\mu \sigma - U(\sigma) \\ & - \frac{1}{4} \Omega^{\mu\nu} \Omega_{\mu\nu} + \frac{1}{2} m_\omega^2 \omega^\mu \omega_\mu - \frac{1}{4} G^{\mu\nu} G_{\mu\nu} + \frac{1}{2} m_\rho^2 \rho^\mu \rho_\mu \\ & - \frac{1}{4} F^{\mu\nu} F_{\mu\nu}. \end{aligned} \quad (3)$$

The strength tensors for the vector mesons and electromagnetic field are defined as $\Omega_{\mu\nu} = \partial_\mu \omega_\nu - \partial_\nu \omega_\mu$, $G_{\mu\nu} = \partial_\mu \rho_\nu - \partial_\nu \rho_\mu$, and $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$. The quark mass m_q ($q = u, d$), characterized by its density dependence, is given by equation (1). The self-interaction potential of the σ field is expressed as

$$U(\sigma) = \frac{1}{2}m_\sigma^2\sigma^2 + \frac{1}{3}b\sigma^3 + \frac{1}{4}c\sigma^4 + B. \quad (4)$$

The bag constant B is introduced such that $U(\sigma_v) = 0$, where σ_v represents the value of the sigma field at which the potential $U(\sigma)$ attains its absolute minimum. The effective quark mass m_q^* is given by

$$m_q^* = m_q - g_\sigma^q \sigma. \quad (5)$$

In nuclear matter, three quarks form a soliton bag, and the effective nucleon mass is derived from the bag energy and expressed as

$$M_N^* = \sum_q E_q = \sum_q \frac{4}{3}\pi R^3 \frac{\Gamma_q}{(2\pi)^3} \int_0^{K_F^q} \sqrt{m_q^{*2} + k^2} \left(\frac{dN_q}{dk} \right) dk, \quad (6)$$

where K_F^q represents the Fermi momentum of the quarks and Γ_q represents the quark degeneracy. dN_q/dk represents the density of states for various types of quarks in a confined space. The expressions for dN_q/dk and K_F^q used in this work can be found in Ref. [106]. The bag radius R can be obtained from the equilibrium condition for the nucleon bag

$$\frac{\delta M_N^*}{\delta R} = 0. \quad (7)$$

We now focus on examining finite nuclei within the IQMDD model. Analogous to the QMC model [74], one can formulate a relativistic Lagrangian density at the hadronic level as

$$\begin{aligned} \mathcal{L}_H = & \bar{\psi} \left[i\gamma^\mu \partial_\mu - M_N^*(\sigma) - g_\omega \gamma^\mu \omega_\mu - \frac{g_\rho}{2} \gamma^\mu \tau \cdot \rho_\mu \right. \\ & - \frac{e}{2} \gamma^\mu (1 + \tau_3) A_\mu + \frac{f_\omega g_\omega}{2M_N} \sigma^{\mu\nu} \partial_\nu \omega_\mu \left. \right] \psi \\ & + \frac{1}{2} \partial^\mu \sigma \partial_\mu \sigma - U(\sigma) - \frac{1}{4} \Omega^{\mu\nu} \Omega_{\mu\nu} \\ & + \frac{1}{2} m_\omega^2 \omega^\mu \omega_\mu - \frac{1}{4} G^{\mu\nu} G_{\mu\nu} + \frac{1}{2} m_\rho^2 \rho^\mu \rho_\mu \\ & - \frac{1}{4} F^{\mu\nu} F_{\mu\nu}. \end{aligned} \quad (8)$$

Starting from the Lagrangian density (8), the RHB equation for the nucleons obtained through a self-consistent treatment of the mean field and pairing correlations is presented as [107]

$$\begin{pmatrix} h_D - \lambda & \Delta \\ -\Delta^* & -h_D^* + \lambda \end{pmatrix} \begin{pmatrix} U_k \\ V_k \end{pmatrix} = E_k \begin{pmatrix} U_k \\ V_k \end{pmatrix}. \quad (9)$$

Here, λ represents the chemical potential, and E_k represents the quasiparticle energy. The single-nucleon Dirac Hamiltonian h_D is given by

$$h_D(\mathbf{r}) = \boldsymbol{\alpha} \cdot \mathbf{p} + V(\mathbf{r}) + \beta M_N^*. \quad (10)$$

The vector potential reads

$$\begin{aligned} V(\mathbf{r}) = & g_\omega \omega_0(\mathbf{r}) \mathbf{I} + \frac{g_\rho}{2} \tau_3 \rho_0(\mathbf{r}) \mathbf{I} - \frac{f_\omega g_\omega}{2M_N} i\gamma^i \partial_i \omega_0(\mathbf{r}) \\ & + \frac{e}{2} (1 + \tau_3) A_0(\mathbf{r}) \mathbf{I}, \end{aligned} \quad (11)$$

where $\mathbf{I}$ represents the identity matrix. Formally, the equations of motion for the mesons and photon are given by

$$\begin{aligned} (-\Delta + m_\sigma^2)\sigma(\mathbf{r}) = & -\frac{\delta M_N^*}{\delta \sigma} \rho_s(\mathbf{r}) - b\sigma^2(\mathbf{r}) - c\sigma^3(\mathbf{r}), \\ (-\Delta + m_\omega^2)\omega_0(\mathbf{r}) = & g_\omega \rho_v(\mathbf{r}) + \frac{f_\omega g_\omega}{2M_N} \rho_0^T(\mathbf{r}), \\ (-\Delta + m_\rho^2)\rho_0(\mathbf{r}) = & \frac{g_\rho}{2} \rho_3(\mathbf{r}), \\ -\Delta A_0(\mathbf{r}) = & e\rho_p(\mathbf{r}). \end{aligned} \quad (12)$$

Here, ρ_s , ρ_v , ρ_0^T , ρ_3 , and ρ_p represent the densities of scalar, vector, tensor, the third component of isovector, and proton, respectively. The relativistic pairing field Δ in Eq. (9) can be expressed as

$$\Delta_{n_1 n'_1} = \frac{1}{2} \sum_{n_2 n'_2} \langle n_1 n'_1 | V^{pp} | n_2 n'_2 \rangle \kappa_{n_2 n'_2}. \quad (13)$$

The pairing tensor is written as

$$\kappa_{nn'} = \sum_{E_k > 0} V_{nk}^* U_{n'k}. \quad (14)$$

In the two-body pairing channel, the finite range force is used, and its form in coordinate space is given by [108]

$$\begin{aligned} V^{pp}(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}'_1, \mathbf{r}'_2) = & -G\delta(\mathbf{R} - \mathbf{R}')P(\mathbf{r})P(\mathbf{r}'), \\ P(\mathbf{r}) = & \frac{1}{(4\pi a^2)^{3/2}} e^{-r^2/2a^2}, \end{aligned} \quad (15)$$

where $\mathbf{r}$ and $\mathbf{R}$ represent the relative and center-of-mass coordinates of the two-nucleon system, respectively. The values of parameters G and a are adopted from Ref. [108].

The aforementioned coupled equations can be solved through a self-consistent iterative scheme, where the wave functions are expanded in terms of the harmonic oscillator basis functions. In our calculation, the number of oscillator shells is set as $N_f = 12$ and $N_b = 20$. From this solution, we can calculate the total binding energy as

$$\begin{aligned}
 B(A, Z) = & \text{Tr}[h_D \rho] + E_{\text{pair}} + E_{\text{c.m.}} - AM_N \\
 & - \frac{1}{2} \int d^3 \mathbf{r} \frac{\partial M_N^*}{\partial \sigma} \rho_s(\mathbf{r}) \sigma(\mathbf{r}) - \frac{1}{6} \int d^3 \mathbf{r} b \sigma^3(\mathbf{r}) \\
 & - \frac{1}{4} \int d^3 \mathbf{r} c \sigma^4(\mathbf{r}) - \frac{1}{2} \int d^3 \mathbf{r} g_{\omega} \rho_v(\mathbf{r}) \omega_0(\mathbf{r}) \\
 & - \frac{1}{2} \int d^3 \mathbf{r} \left[\frac{g_{\rho}}{2} \rho_3(\mathbf{r}) \rho_0(\mathbf{r}) + e \rho_p(\mathbf{r}) A_0(\mathbf{r}) \right] \\
 & - \frac{1}{2} \int d^3 \mathbf{r} \frac{f_{\omega} g_{\omega}}{2M_N} \rho_0^T(\mathbf{r}) \omega_0(\mathbf{r}). \quad (16)
 \end{aligned}$$

Here, E_{pair} represents the nuclear ground-state pairing energy given by

$$E_{\text{pair}} = \frac{1}{4} \sum_{n_1 n'_1} \sum_{n_2 n'_2} \kappa_{n_1 n'_1}^* \langle n_1 n'_1 | V^{pp} | n_2 n'_2 \rangle \kappa_{n_2 n'_2}. \quad (17)$$

The microscopic center-of-mass correction employed in our calculations is given as

$$E_{\text{c.m.}} = -\frac{P_{\text{c.m.}}^2}{2AM_N}, \quad (18)$$

where $P_{\text{c.m.}}$ represents the total momentum as viewed from the center-of-mass frame.

III. NUMERICAL RESULTS AND DISCUSSIONS

Prior to conducting the numerical calculations, we first examine the parameters of this model. The parameter set IQMDD-2*, which incorporates the quark- ω tensor coupling and non-linear ω - ρ interaction term, has been implemented to investigate the properties of finite nuclear systems [109]. This parameter set was initially proposed a decade ago, and subsequent experimental ad-

vancements yielded substantial new data on nuclear mass and neutron skin thickness measurements. To explore potential improvements in finite nuclear property predictions, we conduct a global fitting procedure that combines the simulated annealing algorithm with the Levenberg-Marquardt method, thereby resulting in the parameterization referred to as IQMDD3. The fitted nuclear properties include the binding energies and charge radii of twenty even-even nuclei: ^{16}O , ^{40}Ca , ^{48}Ca , ^{58}Ni , ^{86}Kr , ^{88}Sr , ^{90}Zr , ^{92}Mo , ^{116}Sn , ^{124}Sn , ^{132}Sn , ^{134}Te , ^{138}Ba , ^{142}Nd , ^{144}Sm , ^{148}Gd , ^{186}Pb , ^{208}Pb , ^{210}Pb , and ^{210}Po . Spherical RHB calculations are employed in the fitting process. The experimental binding energy values are adopted from the atomic mass evaluation [110] and charge radii from Ref. [111].

Interaction strengths between the nucleon and ω - and ρ -mesons are selected as $g_{\omega} = 3g_{\omega}^q$ and $g_{\rho} = g_{\rho}^q$ [71]. The bag constant is set to $B = 174.0 \text{ MeV} \cdot \text{fm}^{-3}$. The optimized parameters obtained for the IQMDD3 Lagrangian parameterization are presented in Table 1, with comparative results from the previous IQMDD-2*. To achieve better agreement with experimental neutron skin thickness measurements of ^{208}Pb , the ω - ρ coupling constant (η) is set to zero in the IQMDD3 parameterization. Recent parity-violating electron scattering experiments at JLab have achieved a precision measurement of the neutron skin thickness in ^{208}Pb , which yields a result of $R_n - R_p = 0.283 \pm 0.071 \text{ fm}$ [112]. Using dispersive optical model analysis, Ref. [113] determined the neutron skin thickness of ^{208}Pb to be $0.12 - 0.25 \text{ fm}$. Additional experimental results for this quantity are presented in Fig. 12 of Ref. [114]. The calculated value for ^{208}Pb from the IQMDD3 model is $R_n - R_p = 0.24 \text{ fm}$, which falls within the experimental uncertainty range reported in Refs. [112, 113]. In Table 2, we present IQMDD3 predictions for the total binding energies, charge radii, and neutron-skin thicknesses of the nuclei included in the fit, along with the calculated results for other isotopic chains. All calculations are performed within a spherical RHB approach. For the 20 nuclei used in the fitting, the root-mean-square (rms) deviations from IQMDD3 are 0.023 MeV for the binding energy per nucleon, 2.33 MeV for the total binding energy, and 0.01 fm for the charge radius.

Table 3 presents the spin-orbit splittings for ^{16}O , ^{40}Ca , ^{48}Ca , ^{56}Ni , ^{78}Ni , ^{100}Sn , ^{132}Sn , and ^{208}Pb calculated with the DD-ME2, IQMDD-2*, and IQMDD3 parameter sets. The inclusion of tensor coupling significantly affects nuclear

Table 1. Model parameters of the effective interactions IQMDD3 and IQMDD-2*. Parameter b and the masses of mesons m_{σ} , m_{ω} , and m_{ρ} are measured in MeV.

	m_{σ}	m_{ω}	m_{ρ}	g_{σ}^q	g_{ω}^q	g_{ρ}^q	b	f_{ω}	η
IQMDD-2*	508.0	782.5	763.0	5.09	2.9796	9.7001	-3400.0	2.1	0.03
IQMDD3	498.5998	782.5	763.0	4.9667	2.9404	9.2314	-3114.1540	2.0993	0.0

Table 2. Total binding energies B.E., charge radii R_c , and neutron-skin thickness R_{np} ($R_{np}=R_n-R_p$) calculated with IQM-DD3 compared to the experimental data [110, 111, 115] (values in parentheses).

Nucleus	B.E. /MeV	R_c /fm	R_{np} /fm
¹⁶ O	128.328 (127.619)	2.722 (2.699)	-0.03
⁴⁰ Ca	341.070 (342.052)	3.464 (3.478)	-0.05
⁴⁸ Ca	416.619 (416.001)	3.493 (3.477)	0.20
⁵⁰ Ti	438.112 (437.786)	3.574 (3.570)	0.13
⁵² Cr	455.211 (456.352)	3.656 (3.645)	0.06
⁵⁴ Fe	469.653 (471.765)	3.704 (3.693)	0.00
⁵⁸ Ni	504.782 (506.460)	3.782 (3.776)	0.00
⁶⁸ Ni	590.678 (590.408)	3.899 (3.892 ^a)	0.21
⁸⁶ Kr	751.827 (749.235)	4.189 (4.184)	0.17
⁸⁸ Sr	770.741 (768.468)	4.235 (4.224)	0.13
⁹⁰ Zr	786.180 (783.897)	4.274 (4.269)	0.10
⁹² Mo	799.110 (796.511)	4.319 (4.315)	0.06
¹¹² Sn	953.695 (953.525)	4.583 (4.595)	0.09
¹¹⁶ Sn	989.254 (988.682)	4.616 (4.625)	0.14
¹²⁴ Sn	1051.822 (1049.958)	4.671 (4.674)	0.22
¹³² Sn	1104.904 (1102.843)	4.728 (4.732)	0.30
¹³⁴ Te	1126.579 (1123.408)	4.765 (4.757)	0.26
¹³⁶ Xe	1145.636 (1141.882)	4.805 (4.796)	0.23
¹³⁸ Ba	1162.421 (1158.292)	4.843 (4.838)	0.20
¹⁴⁰ Ce	1176.638 (1172.683)	4.880 (4.877)	0.17
¹⁴² Nd	1188.969 (1185.136)	4.915 (4.912)	0.14
¹⁴⁴ Sm	1199.256 (1195.730)	4.953 (4.952)	0.12
¹⁴⁸ Gd	1222.693 (1220.754)	5.003 (5.008)	0.11
¹⁸⁴ Pb	1435.552 (1432.022)	5.386 (5.393)	0.06
¹⁸⁶ Pb	1454.809 (1451.793)	5.397 (5.403)	0.08
²⁰⁴ Pb	1608.613 (1607.506)	5.491 (5.480)	0.20
²⁰⁸ Pb	1636.544 (1636.430)	5.512 (5.501)	0.24
²¹⁰ Pb	1645.453 (1645.553)	5.528 (5.521)	0.25
²¹⁴ Pb	1661.461 (1663.293)	5.562 (5.558)	0.27
²¹⁰ Po	1647.514 (1645.213)	5.548 (5.570)	0.21

^aAvailable experimental data taken from Ref. [115]; the remaining experimental charge radii are from Ref. [111].

spin-orbit splittings described by the IQMDD model. As the tensor coupling strengths (g_ω^q and f_ω) of the IQMDD3 parameter set are slightly lower than those of IQMDD-2*, the calculated spin-orbit splittings for IQMDD3 are smaller than those for IQMDD-2*, as indicated in Table 3.

In addition, we present the findings of the bulk characteristics for infinite nuclear matter in Table 4. Corrected predictions for nuclear matter properties based on the IQMDD-2* parameter set reported in Ref. [109] are

Table 3. Spin-orbit splittings (in MeV) of proton (π) and neutron (ν) levels in closed-shell nuclei: Calculations with DD-ME2 [61], IQMDD-2*, and IQMDD3 models compared to experimental data from Ref. [116].

Nucleus	State	DD-ME2	IQMDD-2*	IQMDD3	Expt.
¹⁶ O	$\pi(1p)$	6.48	5.81	5.52	6.32
	$\nu(1p)$	6.55	5.86	5.56	6.18
⁴⁰ Ca	$\pi(1d)$	6.70	6.26	5.97	5.40
	$\nu(1d)$	6.77	6.29	6.00	5.63
⁴⁸ Ca	$\pi(2p)$	1.56	2.25	2.16	1.50
	$\nu(2p)$	1.51	1.65	1.57	2.03
⁵⁶ Ni	$\pi(2p)$	1.24	1.60	1.53	1.11
	$\nu(2p)$	1.39	1.78	1.70	1.11
⁷⁸ Ni	$\pi(2p)$	1.31	1.72	1.63	1.40
	$\nu(2p)$	1.36	1.60	1.50	1.33
¹⁰⁰ Sn	$\pi(2p)$	1.41	1.61	1.54	2.85
	$\nu(2d)$	1.95	2.31	2.21	1.65
¹³² Sn	$\pi(2d)$	1.84	2.10	1.99	1.74
	$\nu(2d)$	1.92	2.08	1.96	1.65
²⁰⁸ Pb	$\pi(2d)$	1.68	1.73	1.65	1.34
	$\nu(2g)$	2.34	2.47	2.34	2.49

presented in this table. Compared to IQMDD-2*, the IQMDD3 parameterization, which lacks isoscalar-isovector coupling, yields a larger symmetry energy at saturation density. For the six parameterizations in Table 4, the neutron-skin thickness of ²⁰⁸Pb increases monotonically from 0.160 to 0.285 fm with an increase in the symmetry energy slope from 47.2 to 122.6 MeV. This positive correlation, previously highlighted in the literature [117], indicates that a larger symmetry energy slope L implies higher neutron matter pressure and a thicker neutron-skin thickness in ²⁰⁸Pb.

A. Binding energies

We conduct systematic calculations for even-even isotopes from oxygen ($Z=8$) to oganesson ($Z=118$), focusing on nuclei with available experimental data. Differences in the total binding energy between the calculated and experimental values for even-even nuclei spanning $Z=8-118$ are systematically presented in Fig. 1. The relevant experimental data are obtained from AME2020 atomic mass evaluation [110]. For the specified mass region, our calculations encompass all 868 even-even nuclei in AME2020, containing 646 experimentally measured and 222 estimated mass values. Binding energies and all subsequent results are obtained from axially deformed RHB calculations. As shown in Fig. 1, IQMDD3 and IQMDD-2* exhibit similar overall performance for the total binding energy. In the fitting process, the bind-

Table 4. Bulk properties for infinite nuclear matter at saturation density ρ_0 (in fm^{-3}) across different models [10, 13–15]. Quantities ε_0 , S , L , and K (all in MeV) represent the binding energy per nucleon, symmetry energy, slope of the symmetry energy, and incompressibility coefficient of symmetric nuclear matter at saturation density. The neutron-skin thickness of ^{208}Pb represented by R_{np} (in fm) is also included.

Model	ρ_0	ε_0	S	L	K	R_{np}
IU-FSU	0.155	-16.40	31.30	47.2	231.3	0.160
FSU	0.148	-16.30	32.59	60.5	230.0	0.205
IQMDD-2*	0.155	-16.38	34.16	72.8	283.3	0.212
IQMDD3	0.153	-16.41	36.00	100.1	292.8	0.236
NL3	0.148	-16.24	37.40	118.5	271.8	0.277
NL3*	0.150	-16.31	38.68	122.6	258.1	0.285

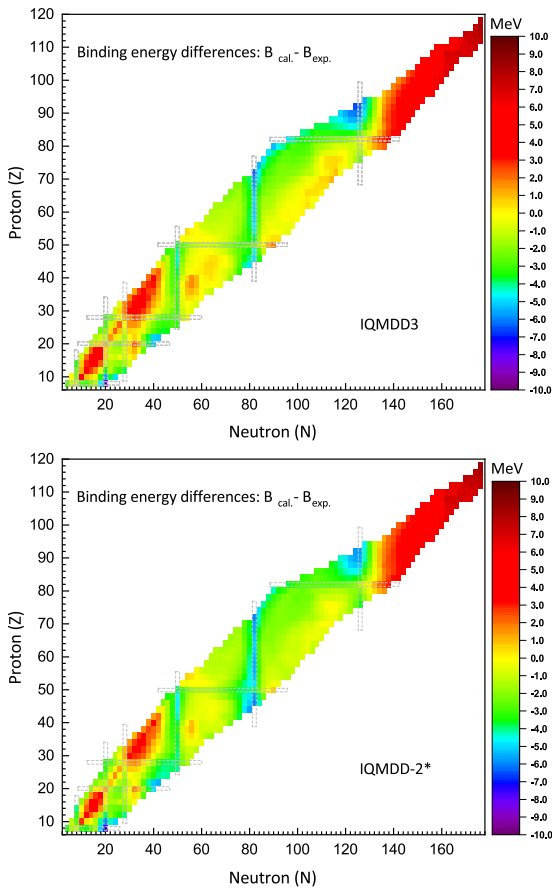


Fig. 1. (color online) Differences in total binding energy between the calculations and experimental data for even-even nuclei using the IQMDD3 and IQMDD-2*. To guide the eye, the dashed frames correspond to the positions of the magic numbers 8, 20, 28, 50, 82, and 126.

ing energy per nucleon is employed as a constraint instead of the total binding energy. This adjustment effectively increased the fitting weight for light nuclei but led to a moderately larger deviation in the total binding energy of heavy nuclei compared to that of experimental values. This is reflected in the relatively large deviation

in the total binding energy for heavy nuclei shown in Fig. 1. Statistical analysis reveals that the rms deviation of binding energy per nucleon is $\sigma_{B/A} = 0.034$ MeV for IQMDD3 and $\sigma_{B/A} = 0.036$ MeV for IQMDD-2*. For the total binding energy, the corresponding values are $\sigma_B = 2.893$ MeV for IQMDD3 and $\sigma_B = 2.984$ MeV for IQMDD-2*. In addition, the statistical analysis shows that both models perform better in the medium and heavy nucleus regions compared to the light nucleus region in terms of the average binding energy per nucleon. When the analysis is confined to the range from calcium ($Z=20$) to oganesson ($Z=118$), the rms deviation decreases significantly to $\sigma_{B/A} = 0.020$ MeV for IQMDD3 and $\sigma_{B/A} = 0.022$ MeV for IQMDD-2*.

For comparison, the typical binding energy rms deviations of established EDFs are 1.518 MeV for PC-PK1 when including the rotational correction energies [118], 1.74 MeV for QMC π -III-T [119], 2.96 MeV for NL3* [120], 2.39 MeV for DD-ME2 [120], 2.29 MeV for DD-ME δ [120], 2.01 MeV for DD-PC1 [120], and 1.339 MeV for 91 spherical nuclei for PC-L3R [121]. In terms of predicting nuclear binding energies, the performance of IQMDD3 is comparable overall to that of NL3*.

B. Quadrupole deformations

Figure 2 illustrates the theoretical predictions of ground-state quadrupole deformation parameters β_2 for all isotopic chains examined in this work. The deformation parameter β_2 is derived from the calculated nucleon quadrupole moments Q via [7].

$$Q = \sqrt{\frac{16\pi}{5}} \frac{3}{4\pi} AR_0^2 \beta_2, \quad (19)$$

with $R_0 = 1.2A^{1/3}$ (fm). The β_2 parameter, a dimensionless quantity with an absolute value less than 1, is characterized by relatively large empirical uncertainties. Given that our primary objective is to investigate the systematic trends and global characteristics of nuclear deformation patterns predicted by the IQMDD3 model, a comparison

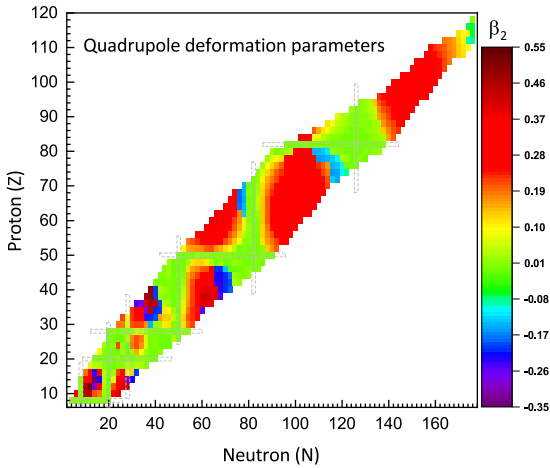


Fig. 2. (color online) Quadrupole deformation parameters for even-even nuclei with proton numbers $8 \leq Z \leq 118$ calculated using the RHB approach with the IQMDD3 interaction.

with experimental values is not included in the present analysis. As illustrated in Fig. 2, nuclei with spherical configurations are predominantly located in the region around magic numbers. Nuclei with proton and neutron numbers situated between two magic numbers exhibit larger deformations. In addition, for nuclei located in the middle of the major shells, deformations in the light and medium-mass regions are more pronounced than those in heavy and superheavy nuclei.

For the neutron magic numbers $N=8, 20, 50, 82$, and 126 , the corresponding isotonic chains are predominantly composed of spherical nuclei. However, an isotonic chain with neutron number $N=28$ exhibits significant shape shifts. Isotopic chains characterized by proton magic numbers ($Z=8, 20, 50, 82$) maintain spherical configurations, whereas the $Z=28$ isotopic chain exhibits moderate quadrupole deformation, which is not as pronounced as that in the isotonic chain with $N=28$.

For nuclei with neutron numbers $Z > 50$, those located in the central region between major nuclear shells predominantly exhibit prolate deformation. A characteristic shape transition occurs as the neutron number approaches magic numbers, where nuclear deformation evolves from prolate to oblate configurations. Within the proton number range of $28 < Z < 50$, oblate nuclear configurations are predominantly found near the central region of the major shells. Similarly, the range of $8 < Z < 20$ contains oblate nuclei. However, according to the IQMDD3 model, no significantly deformed oblate nuclei are predicted to exist within the $20 < Z < 28$ range.

C. Charge radii

Besides nuclear deformations, the charge radius is also a crucial quantity for characterizing finite nuclei and is expressed as

$$R_c = \sqrt{R_p^2 + 0.64} \quad (\text{fm}), \quad (20)$$

where the factor 0.64 corrects for the finite size of the proton charge distribution. The deviations between the theoretical values derived from the IQMDD3 model and experimental data cited in Refs. [111, 115] are depicted in Fig. 3. IQMDD3 predictions are in reasonable agreement with the experimental charge radii, with calculated deviations within the ± 0.04 fm range. For the 368 even-even nuclei with known charge radii, the rms deviation between the theoretical and experimental values is $\sigma = 0.022$ fm compared to $\sigma = 0.021$ fm for IQMDD-2*.

The most significant discrepancies occur in the Mg ($Z=12$), Si ($Z=14$), and Cm ($Z=96$) isotopic chains, where theoretical values exceed experimental values by over 0.05 fm. For ^{26}Mg , ^{28}Si , and ^{30}Si , the theoretical overestimation of the charge radii relative to experimental values can be attributed to the significant oblate deformation ($|\beta_2| > 0.2$) predicted in their ground states. Experiments reveal that the Cm ($Z=96$) isotopic chain exhibits anomalous charge radii, which are smaller than those of Pu ($Z=94$) [111]. However, IQMDD3 calculations indicate that charge radii increase with proton number in an isotonic chain. This results in a noticeable discrepancy between theoretical predictions and experimental observations for the charge radii of the Cm isotopic chain. This anomalous charge radius behavior exhibited by U-Pu-Cm isotopes has already been extensively discussed in Refs. [118, 120, 122].

D. Two-neutron and two-proton separation energies

Two-nucleon separation energies (S_{2n} and S_{2p}) offer essential insights into the investigation of neutron and proton drip lines, nuclear shell evolution, and nucleon ra-

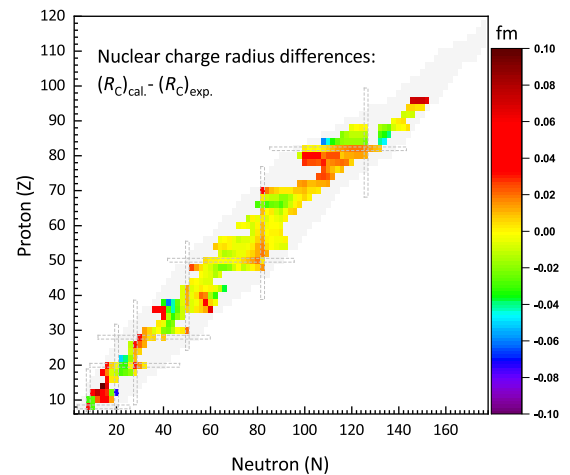


Fig. 3. (color online) Discrepancies between IQMDD3+RHB calculations and experimental data for even-even nuclei with measured charge radii.

dioactivity phenomena [123]. They can be expressed as

$$\begin{aligned} S_{2n}(A, Z) &= B(A, Z) - B(A - 2, Z), \\ S_{2p}(A, Z) &= B(A, Z) - B(A - 2, Z - 2), \end{aligned} \quad (21)$$

where $B(A, Z)$ represents the binding energy of a nucleus with mass number A and proton number Z , calculated from equation (16). In Fig. 4, the predicted two-neutron separation energies S_{2n} for even-even nuclei using the IQMDD3 parameters are depicted along with the discrepancies between the theoretical predictions and experimental data from Ref. [110]. The left panel of Fig. 4 shows that, for a given isotopic chain, S_{2n} decreases with increasing neutron number, while for a given isotonic chain, S_{2n} increases with the increase in proton number. Meanwhile, the two-neutron separation energy S_{2n} on the proton-rich side decreases gradually with an increase in the proton number. The theoretical values of S_{2n} calculated by IQMDD3 are compared with the experimental values. As shown in the right panel of Fig. 4, most differences between theory and experiment remain within ± 2 MeV, and for the studied nuclei, the rms of the differences is determined to be 0.984 MeV.

As seen from the left panel of Fig. 4, there are pronounced sudden drops in S_{2n} at the positions of $N=50, 82$, and 126 . Similar sudden drops can also be observed at $N=20$ and 28 . These abrupt decreases in S_{2n} indicate the emergence of neutron shells, which are well-reproduced by IQMDD3.

Figure 5 presents the systematic trends of two-proton separation energy (S_{2p}) for even-even nuclei, accompanied by the differences between theoretical predictions and experimental measurements. In contrast to the two-neutron separation energy (S_{2n}), for a given isotopic chain, S_{2p} exhibits an increasing trend with neutron number, while in isotonic chains, it demonstrates a decreasing

trend with proton number. Furthermore, the data indicate a gradual reduction in S_{2p} values within neutron-rich regions with an increase in the proton number. The right panel of Fig. 5 demonstrates that the majority of theoretical-experimental deviations fall within a range of ± 2 MeV. For the nuclei under investigation, the rms deviation of S_{2p} between theory and experiment is calculated to be 0.848 MeV, indicating acceptable agreement between theoretical predictions and experimental observations.

As observed in the left panel of Fig. 5, there are notable sudden drops in S_{2p} at the proton numbers $Z=20, 28$, and 50 . These abrupt decreases in S_{2p} signify the appearance of proton shells. For the isotope chain with $Z=82$, the variation in S_{2p} is relatively subtle. However, the two-proton shell gaps, which will be discussed in the following subsection, provide a clearer manifestation of this change.

E. Two-nucleon shell gaps

The two-nucleon energy gaps essential for examining the shell effects in finite nuclei are defined as

$$\begin{aligned} \delta_{2n}(A, Z) &= S_{2n}(A, Z) - S_{2n}(A + 2, Z), \\ \delta_{2p}(A, Z) &= S_{2p}(A, Z) - S_{2p}(A + 2, Z + 2). \end{aligned} \quad (22)$$

Figure 6 displays the two-neutron gaps (δ_{2n}) for even-even nuclei in the region of $8 \leq Z \leq 114$. δ_{2n} exhibits significant sharp changes at neutron numbers $N=8, 20, 28, 50, 82$, and 126 , which can be considered one of the signatures of magic numbers. For the magic number $N=28$, δ_{2n} becomes relatively small in both proton-rich and neutron-rich regions, which suggests that the $N=28$ shell closure may be quenched in these areas.

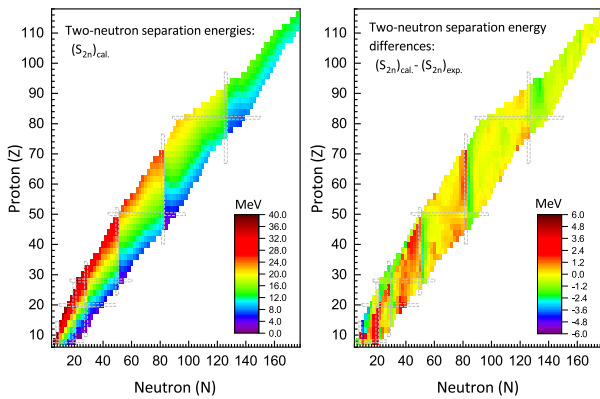


Fig. 4. (color online) Left: The two-neutron separation energies of even-even nuclei calculated by IQMDD3. Right: Deviation between the theoretical and experimental values of the two-neutron separation energies.

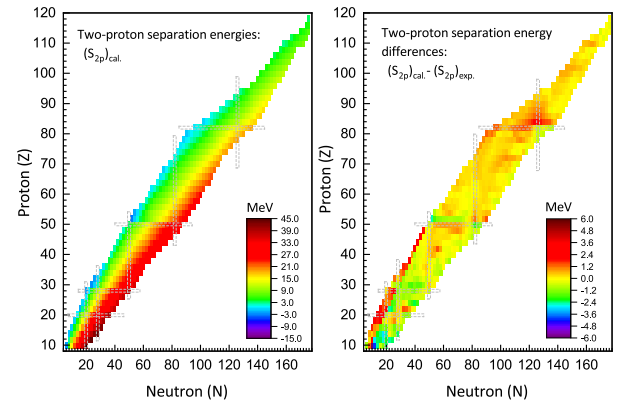


Fig. 5. (color online) Left: Theoretical predictions of two-proton separation energies for even-even nuclei obtained from IQMDD3 calculations. Right: Discrepancy between theoretical calculations and experimental measurements of the two-proton separation energies.

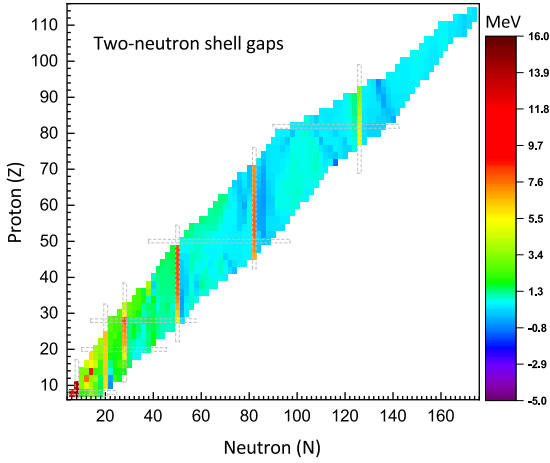


Fig. 6. (color online) Two-neutron shell gaps δ_{2n} for even-even nuclei, as calculated with the IQMDD3 parameterization.

Figure 7 presents the two-proton gaps (δ_{2p}) for even-even nuclei in the region of $10 \leq Z \leq 116$. δ_{2p} exhibits distinct peaks at proton numbers $Z=20, 28$, and 50 , which is consistent with conclusions regarding shell closures shown in Fig. 5. For $Z=82$, δ_{2p} also shows a significant sharp change, whereas the change in S_{2p} is less evident in Fig. 5. This implies that, compared to the two-proton separation energy, the two-proton gap provides a more intuitive indication of potential magic numbers. In addition, it is noteworthy that a relatively large δ_{2p} is observed at $Z=14$, suggesting that $Z=14$ may correspond to a subshell closure.

F. α -decay energies

α -decay usually occurs in heavy and superheavy radioactive nuclides. The α -decay energy, denoted as Q_α , is expressed as

$$Q_\alpha = B(A-4, Z-2) + B(2, 2) - B(A, Z). \quad (23)$$

Figure 8 presents discrepancies between Q_α values calculated using the IQMDD3 model and experimental data from Ref. [110]. Among the 405 studied even-even nuclei with positive Q_α values, relatively larger deviations are observed in specific nuclei of the $Z=52$ and $Z=94$ isotopic chains, as well as in some nuclei along the $N=84$ isotonic chain. Overall, most discrepancies are within ± 1.5 MeV, and the rms of the discrepancies is 0.733 MeV. Recently, machine learning approaches have been successfully employed in the description of α -decay energies, achieving a remarkably high level of precision [124, 125].

IV. SUMMARY

In this study, we developed an RHB model incorporating quark-meson couplings and established a new para-

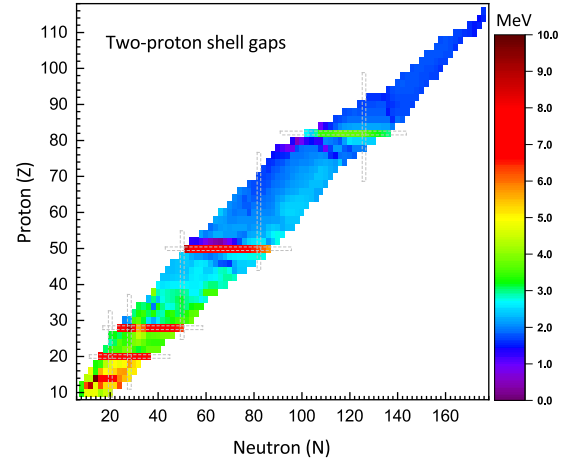


Fig. 7. (color online) Two-proton shell gaps δ_{2p} in even-even nuclei from IQMDD3 calculations.

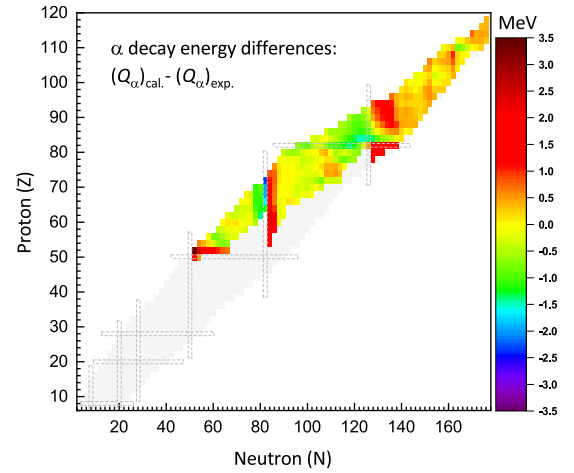


Fig. 8. (color online) Discrepancies between the IQMDD3-calculated and experimental values of α decay energy Q_α .

meter set named IQMDD3 by fitting to the binding energies and charge radii of 20 selected even-even nuclei. Using this parameter set, we systematically studied the ground-state properties of 868 even-even nuclei, with proton numbers from $Z=8$ to $Z=118$. For the twenty nuclei employed for fitting, the rms deviation of the binding energy per nucleon obtained with the IQMDD3 model was 0.023 MeV, that of the total binding energy was 2.33 MeV, and that of the charge radius was 0.01 fm. The predicted neutron skin thickness for ^{208}Pb with this parameter set was 0.24 fm, with the symmetry energy and its slope at saturation density were 36.00 and 100.1 MeV, respectively.

For the 868 even-even nuclei studied, the rms deviation of the binding energy per nucleon was 0.034 MeV, and the rms deviation of the total binding energy was 2.893 MeV. We investigated systematic trends and general features of quadrupole deformations in these even-even nuclei. Results showed that nuclei in the magic number regions and those in their vicinity were mostly spherical,

except for the $N=28$ isotonic chain, which exhibited significant shape shifts. In contrast, nuclei far from the major shells exhibited larger deformations, with the largest deformations often occurring in the central regions between major shells. The analysis of charge radii showed that, for the 368 studied even-even nuclei with available experimental data, the rms deviation of the charge radii was 0.022 fm.

We systematically examined the two-nucleon separation energies, as well as the two-nucleon shell gaps. For the same isotopic chain, the two-neutron separation energy was larger in the proton-rich regions than in the neutron-rich regions, while the two-proton separation energy was the opposite, being larger in neutron-rich re-

gions than that in proton-rich regions. The conventional magic numbers were successfully reconstructed based on the systematic trends observed in two-nucleon separation energies and two-nucleon shell gaps. In addition, the two-proton shell gaps indicated that $Z = 14$ might be a subshell closure.

Finally, for nuclei with positive α -decay energies, the extracted Q_α values were compared with the 405 available experimental data, yielding an rms deviation of $\sigma = 0.733$ MeV.

To further examine the validity and reliability of the IQMDD3 parameter set in describing finite nuclei, future work will employ this model to study the properties of drip-line nuclei far from the β -stability line.

References

- [1] D. Vautherin and D. M. Brink, *Phys. Rev. C* **5**, 626 (1972)
- [2] D. Vautherin, *Phys. Rev. C* **7**, 296 (1973)
- [3] J. Dechargé and D. Gogny, *Phys. Rev. C* **21**, 1568 (1980)
- [4] J. D. Walecka, *Ann. Phys. (NY)* **83**, 491 (1974)
- [5] B. D. Serot and J. D. Walecka, *Adv. Nucl. Phys.* **16**, 1 (1986)
- [6] C. J. Horowitz and B. D. Serot, *Nucl. Phys. A* **368**, 503 (1981)
- [7] Y. K. Gambhir, P. Ring, and A. Thimet, *Ann. Phys. (N.Y.)* **198**, 132 (1990)
- [8] Y. Sugahara and H. Toki, *Nucl. Phys. A* **579**, 557 (1994)
- [9] P. Ring, *Prog. Part. Nucl. Phys.* **37**, 193 (1996)
- [10] G. A. Lalazissis, J. König, and P. Ring, *Phys. Rev. C* **55**, 540 (1997)
- [11] G. A. Lalazissis, S. Raman, and P. Ring, *At. Data. Nucl. Data. Tab.* **71**, 1 (1999)
- [12] M. Bender, P.-H. Heenen, and P.-G. Reinhard, *Rev. Mod. Phys.* **75**, 121 (2003)
- [13] B. G. Todd-Rutel and J. Piekarewicz, *Phys. Rev. Lett.* **95**, 122501 (2005)
- [14] G. A. Lalazissis, S. Karatzikos, R. Fossion *et al.*, *Phys. Lett. B* **671**, 36 (2009)
- [15] F. J. Fattoyev, C. J. Horowitz, J. Piekarewicz *et al.*, *Phys. Rev. C* **82**, 055803 (2010)
- [16] N. Schunck and J. L. Egido, *Phys. Rev. C* **78**, 064305 (2008)
- [17] S. Mizutori, J. Dobaczewski, G. A. Lalazissis *et al.*, *Phys. Rev. C* **61**, 044326 (2000)
- [18] R. J. Furnstahl, *Nucl. Phys. A* **706**, 85 (2002)
- [19] P. Klüpfel, P.-G. Reinhard, T. J. Bürvenich *et al.*, *Phys. Rev. C* **79**, 034310 (2009)
- [20] H. Sagawa, *Phys. Lett. B* **286**, 7 (1992)
- [21] Z. Z. Ren, G. O. Xu, B. Q. Chen *et al.*, *Phys. Lett. B* **351**, 11 (1995)
- [22] Z. Z. Ren, A. Faessler, and A. Bobyk, *Phys. Rev. C* **57**, 2752 (1998)
- [23] Z. Z. Ren, B. Q. Chen, Z. Y. Ma *et al.*, *Phys. Rev. C* **53**, R572 (1996)
- [24] M. Grasso, S. Yoshida, N. Sandulescu *et al.*, *Phys. Rev. C* **74**, 064317 (2006)
- [25] J. C. Pei, F. R. Xu, and P. D. Stevenson, *Nucl. Phys. A* **765**, 29 (2006)
- [26] K. Y. Zhang, P. Papakonstantinou, M.-H. Mun *et al.*, *Phys. Rev. C* **107**, L041303 (2023)
- [27] S. G. Zhou, J. Meng, P. Ring *et al.*, *Phys. Rev. C* **82**, 011301 (2010)
- [28] S. Ćwiok, J. Dobaczewski, P.-H. Heenen *et al.*, *Nucl. Phys. A* **611**, 211 (1996)
- [29] P. Demetriou and S. Goriely, *Nucl. Phys. A* **695**, 95 (2001)
- [30] A. T. Kruppa, M. Bender, W. Nazarewicz *et al.*, *Phys. Rev. C* **61**, 034313 (2000)
- [31] R. Y. Zheng, X. X. Sun, G. F. Shen *et al.*, *Chin. Phys. C* **48**, 014107 (2024)
- [32] P.-G. Reinhard, D. J. Dean, W. Nazarewicz *et al.*, *Phys. Rev. C* **60**, 014316 (1999)
- [33] G. A. Lalazissis, D. Vretenar, P. Ring *et al.*, *Phys. Rev. C* **60**, 014310 (1999)
- [34] F. Sarazin, *et al.*, *Phys. Rev. Lett.* **84**, 5062 (2000)
- [35] Z. Z. Ren, *Phys. Rev. C* **65**, 051304 (2002)
- [36] S. Quan, Q. Chen, Z. P. Li *et al.*, *Phys. Rev. C* **95**, 054321 (2017)
- [37] J. Ljungvall, *et al.*, *Phys. Rev. Lett.* **100**, 102502 (2008)
- [38] B. A. Brown, *Phys. Rev. C* **58**, 220 (1998)
- [39] T. Nikšić, D. Vretenar, and P. Ring, *Prog. Part. Nucl. Phys.* **66**, 519 (2011)
- [40] J. Dobaczewski, W. Nazarewicz, T. R. Werner *et al.*, *Phys. Rev. C* **53**, 2809 (1996)
- [41] S. Goriely, F. Tondeur, and J. M. Pearson, *At. Data Nucl. Data Tables* **77**, 311 (2001)
- [42] M. V. Stoitsov, J. Dobaczewski, W. Nazarewicz *et al.*, *Phys. Rev. C* **68**, 054312 (2003)
- [43] S. Hilaire and M. Girod, *Eur. Phys. J. A* **33**, 237 (2007)
- [44] L. Wang, Q. L. Niu, J. J. Zhang *et al.*, *Sci. China-Phys. Mech. Astron.* **66**, 102011 (2023)
- [45] H. K. Wang, Q. Su, C. Xu *et al.*, *J. Phys. G: Nucl. Part. Phys.* **50**, 095104 (2023)
- [46] H. T. Zhang, C. F. Chen, X. R. Zhou *et al.*, *Chin. Phys. C* **49**, 044101 (2025)
- [47] Z. Q. Sheng, Z. Z. Ren, and W. Z. Jiang, *Nucl. Phys. A* **832**, 49 (2010)
- [48] E. Chabanat, P. Bonche, P. Haensel *et al.*, *Nucl. Phys. A* **627**, 710 (1997)
- [49] H. Shen, H. Toki, K. Oyamatsu, and K. Sumiyoshi, *Nucl. Phys. A* **637**, 435 (1998)
- [50] T. Malik, N. Alam, M. Fortin *et al.*, *Phys. Rev. C* **98**, 035804 (2018)
- [51] W.-C. Chen and J. Piekarewicz, *Phys. Rev. C* **90**, 044305 (2014)
- [52] B. Hong and Z. Z. Ren, *Phys. Rev. D* **109**, 023002 (2024)
- [53] J. Boguta and A. R. Bodmer, *Nucl. Phys. A* **292**, 413

- (1977)
- [54] S. Gmuca, *Nucl. Phys. A* **547**, 447 (1992)
- [55] C. J. Horowitz and J. Piekarewicz, *Phys. Rev. Lett.* **86**, 5647 (2001)
- [56] C. J. Horowitz and J. Piekarewicz, *Phys. Rev. C* **64**, 062802 (2001)
- [57] B. T. Reed, F. J. Fattoyev, C. J. Horowitz *et al.*, *Phys. Rev. C* **109**, 035803 (2024)
- [58] J. Zimanyi and S. A. Moszkowski, *Phys. Rev. C* **42**, 1416 (1990)
- [59] M. Rufa, P.-G. Reinhard, J. A. Maruhn, *et al.*, *Phys. Rev. C* **38**, 390 (1988)
- [60] R. Brockmann and H. Toki, *Phys. Rev. Lett.* **68**, 3408 (1992)
- [61] G. A. Lalazissis, T. Nikšić, D. Vretenar *et al.*, *Phys. Rev. C* **71**, 024312 (2005)
- [62] S. Typel and H. H. Wolter, *Nucl. Phys. A* **656**, 331 (1999)
- [63] W. H. Long, J. Meng, N. V. Giai *et al.*, *Phys. Rev. C* **69**, 034319 (2004)
- [64] T. Nikšić, D. Vretenar, P. Finelli *et al.*, *Phys. Rev. C* **66**, 024306 (2002)
- [65] B. A. Nikolaus, T. Hoch, and D. G. Madland, *Phys. Rev. C* **46**, 1757 (1992)
- [66] J. J. Rusnak and R. J. Furnstahl, *Nucl. Phys. A* **627**, 495 (1997)
- [67] T. Bürvenich, D. G. Madland, J. A. Maruhn *et al.*, *Phys. Rev. C* **65**, 044308 (2002)
- [68] P. W. Zhao, Z. P. Li, J. M. Yao *et al.*, *Phys. Rev. C* **82**, 054319 (2010)
- [69] P. A. M. Guichon, *Phys. Lett. B* **200**, 235 (1988)
- [70] S. Fleck, W. Bentz, K. Shimizu *et al.*, *Nucl. Phys. A* **510**, 731 (1990)
- [71] K. Saito and A. W. Thomas, *Phys. Lett. B* **327**, 9 (1994)
- [72] D. P. Menezes, P. K. Panda, and C. Providência, *Phys. Rev. C* **72**, 035802 (2005)
- [73] P. A. M. Guichon, J. R. Stone, and A. W. Thomas, *Prog. Part. Nucl. Phys.* **100**, 262 (2018)
- [74] K. Saito, K. Tsushima, and A. W. Thomas, *Prog. Part. Nucl. Phys.* **58**, 1 (2007)
- [75] P. Papazoglou, S. Schramm, J. Schaffner-Bielich *et al.*, *Phys. Rev. C* **57**, 2576 (1998)
- [76] P. Papazoglou, D. Zschesche, S. Schramm *et al.*, *Phys. Rev. C* **59**, 411 (1999)
- [77] P. Wang, V. E. Lyubovitskij, T. Gutsche *et al.*, *Phys. Rev. C* **67**, 015210 (2003)
- [78] P. Wang, Z. Y. Zhang, Y. W. Yu *et al.*, *Nucl. Phys. A* **688**, 791 (2001)
- [79] H. Toki, U. Meyer, A. Faessler *et al.*, *Phys. Rev. C* **58**, 3749 (1998)
- [80] H. Shen and H. Toki, *Phys. Rev. C* **61**, 045205 (2000)
- [81] J. N. Hu, A. Li, H. Toki *et al.*, *Phys. Rev. C* **89**, 025802 (2014)
- [82] C. Wu, W.-L. Qian, and R.-K. Su, *Phys. Rev. C* **72**, 035205 (2005)
- [83] C. Wu, W.-L. Qian, and R.-K. Su, *Phys. Rev. C* **77**, 015203 (2008)
- [84] C. Wu and R.-K. Su, *J. Phys. G: Nucl. Part. Phys.* **36**, 095101 (2009)
- [85] G. N. Fowler, S. Raha, and R. M. Weiner, *Z. Phys. C* **9**, 271 (1981)
- [86] G. X. Peng, H. C. Chiang, J. J. Yang *et al.*, *Phys. Rev. C* **61**, 015201 (1999)
- [87] O. G. Benvenuto and G. Lugones, *Phys. Rev. D* **51**, 1989 (1995)
- [88] A. Chodos, R. L. Jaffe, K. Johnson *et al.*, *Phys. Rev. D* **10**, 2599 (1974)
- [89] A. Chodos, R. L. Jaffe, K. Johnson *et al.*, *Phys. Rev. D* **9**, 3471 (1974)
- [90] E. P. Gilson and R. L. Jaffe, *Phys. Rev. Lett.* **71**, 332 (1993)
- [91] C. Wu and Z. Z. Ren, *J. Phys. G: Nucl. Part. Phys.* **37**, 105110 (2010)
- [92] W. L. Zhan, *et al.*, *Nucl. Phys. A* **834**, 694c (2010)
- [93] T. Motobayashi, *Nucl. Phys. A* **834**, 707c (2010)
- [94] M. Thoennessen, *Nucl. Phys. A* **834**, 688c (2010)
- [95] D. Vretenar, A. V. Afanasjev, G. A. Lalazissis *et al.*, *Phys. Rep.* **409**, 101 (2005)
- [96] J. Meng and P. Ring, *Phys. Rev. Lett.* **77**, 3963 (1996)
- [97] W. Pöschl, D. Vretenar, G. A. Lalazissis *et al.*, *Phys. Rev. Lett.* **79**, 3841 (1997)
- [98] Y. Tian, Z. Y. Ma, and P. Ring, *Phys. Lett. B* **676**, 44 (2009)
- [99] J. Geng and W. H. Long, *Phys. Rev. C* **105**, 034329 (2022)
- [100] K. Zhang, *et al.*, *Phys. Rev. C* **102**, 024314 (2020)
- [101] S.-G. Zhou, J. Meng, S. Yamaji *et al.*, *Chin. Phys. Lett.* **17**, 717 (2000)
- [102] Y. N. Zhang, J. C. Pei, and F. R. Xu, *Phys. Rev. C* **88**, 054305 (2013)
- [103] J. Meng, *Nucl. Phys. A* **635**, 3 (1998)
- [104] S.-G. Zhou, J. Meng, and P. Ring, *Phys. Rev. C* **68**, 034323 (2003)
- [105] K. Y. Zhang, C. Pan, and S. Q. Zhang, *Phys. Rev. C* **106**, 024302 (2022)
- [106] Y. Zhang, W. L. Qian, S. Q. Ying *et al.*, *J. Phys. G: Nucl. Part. Phys.* **27**, 2241 (2001)
- [107] H. Kucharek and P. Ring, *Z. Phys. A* **339**, 23 (1991)
- [108] T. Nikšić, N. Paar, D. Vretenar *et al.*, *Comput. Phys. Commun.* **185**, 1808 (2014)
- [109] R. L. Xu, C. Wu, W. L. Qian *et al.*, *Eur. Phys. J. A* **49**, 141 (2013)
- [110] M. Wang, W. J. Huang, F. G. Kondev *et al.*, *Chin. Phys. C* **45**, 030003 (2021)
- [111] I. Angeli and K. P. Marinova, *At. Data. Nucl. Data. Tab.* **99**, 69 (2013)
- [112] D. Adhikari, *et al.*, *Phys. Rev. Lett.* **126**, 172502 (2021)
- [113] C. D. Pruitt, R. J. Charity, L. G. Sobotka *et al.*, *Phys. Rev. Lett.* **125**, 102501 (2020)
- [114] Q. Zhao, Z. X. Ren, P. W. Zhao *et al.*, *Phys. Rev. C* **106**, 034315 (2022)
- [115] T. Li, Y. Luo, and N. Wang, *At. Data. Nucl. Data. Tab.* **140**, 101440 (2021)
- [116] H. Grawe, K. Langanke, and G. Martínez-Pinedo, *Rep. Prog. Phys.* **70**, 1525 (2007)
- [117] X. Roca-Maza, M. Centelles, X. Viñas *et al.*, *Phys. Rev. Lett.* **106**, 252501 (2011)
- [118] K. Zhang, *et al.*, *At. Data. Nucl. Data. Tab.* **144**, 101488 (2022)
- [119] K. L. Martinez, A. W. Thomas, P. A. M. Guichon *et al.*, *Phys. Rev. C* **102**, 034304 (2020)
- [120] S. E. Agbemava, A. V. Afanasjev, D. Ray *et al.*, *Phys. Rev. C* **89**, 054320 (2014)
- [121] Z. X. Liu, Y. H. Lam, N. Lu *et al.*, *Phys. Lett. B* **842**, 137946 (2023)
- [122] P. Guo, *et al.*, *At. Data. Nucl. Data. Tab.* **158**, 101661 (2024)
- [123] H. L. Wang, Z. Wang, D. Bai *et al.*, *Chin. Phys. C* **49**, 044111 (2025)
- [124] Z. Y. Yuan, D. Bai, Z. Z. Ren *et al.*, *Chin. Phys. C* **46**, 024101 (2022)
- [125] Z. Y. Yuan, D. Bai, Z. Wang *et al.*, *Nucl. Sci. Tech.* **35**, 105 (2024)