

# Light higgsino-dominated NLSPs in semi-constrained NMSSM\*

Kun Wang(王坤) Jingya Zhu(朱经亚)<sup>1)</sup>

Center for Theoretical Physics, School of Physics and Technology, Wuhan University, Wuhan 430072, China

**Abstract:** In the semi-constrained next-to minimal supersymmetric standard model (scNMSSM, or NMSSM with non-universal Higgs mass) under current constraints, we consider a scenario where  $h_2$  is the SM-like Higgs,  $\tilde{\chi}_1^0$  is a singlino-dominated LSP;  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_{2,3}^0$  are mass-degenerated, light, and higgsino-dominated next-to-lightest supersymmetric particles (NLSPs). We investigate the constraints of these NLSPs by searching for supersymmetry particles at the LHC Run-I and Run-II and discuss the possibility of discovering these NLSPs in the future. We arrive at the following conclusions: (i) With all data of Run I and up to  $36 \text{ fb}^{-1}$  data of Run II at the LHC, the search results by ATLAS and CMS still cannot exclude the higgsino-dominated NLSPs of  $100 \sim 200 \text{ GeV}$ . (ii) When the mass difference with  $\tilde{\chi}_1^0$  is smaller than  $m_{h_2}$ ,  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_3^0$  have opposite preferences with regard to decaying to  $Z/Z^*$  or  $h_1$ . (iii) When the mass difference between NLSP and LSP is larger than  $m_Z$ , most samples can be verified at the  $5\sigma$  level with future  $300 \text{ fb}^{-1}$  data at the LHC. Meanwhile, with  $3000 \text{ fb}^{-1}$  data at high-luminosity LHC (HL-LHC), almost all of the samples can be verified at the  $5\sigma$  level, even if the mass difference is insufficient. (iv) The  $a_1$  funnel and the  $h_2/Z$  funnel mechanisms for the singlino-dominated LSP annihilation cannot be distinguished by searching for NLSPs.

**Keywords:** NMSSM, supersymmetry phenomenology, higgsino

**DOI:** 10.1088/1674-1137/44/6/061001

## 1 Introduction

As an internal symmetry between fermions and bosons, supersymmetry (SUSY) is an attractive concept. In the framework of SUSY, the strong, weak, and hypercharge gauge couplings ( $g_3, g_2, g_1$ ) can be unified at the GUT scale ( $\sim 10^{16} \text{ GeV}$ ), and the large hierarchy problem between the electroweak and the Planck scales can be resolved. Further, with the  $R$ -parity conserved, the lightest SUSY particle (LSP) is stable and can be a good candidate for weakly-interacting-massive-particle (WIMP) dark matter (DM).

SUSY at the TeV scale is motivated by possible cancellation of quadratic divergences of the Higgs boson mass. The simplest implementation of SUSY is the minimal supersymmetric extension to the standard model (MSSM). Since the soft SUSY breaking parameters are totally free in the MSSM, a dynamic approach to obtain these parameters is more favored. In the minimal super-

gravity (mSUGRA), the Kähler potential is employed to yield minimal kinetic energy terms for MSSM fields, where all trilinear couplings, gaugino, and scalar mass parameters unify respectively at the GUT scale. The fully constrained MSSM (CMSSM) is the MSSM with the boundary conditions that are the same as in the mSUGRA. However, to obtain a 125 GeV SM-like Higgs, the MSSM requires large one-loop radiative corrections to the Higgs mass, which renders the MSSM not natural. There is a so-called  $\mu$ -problem [1] in the MSSM, where the superpotential contains a term  $\mu \hat{H}_u \hat{H}_d$ , and  $\mu$  is the only dimensionful parameter, which has to be chosen artificially.

The next-to minimal supersymmetric standard model (NMSSM) solves the  $\mu$ -problem by introducing a complex singlet superfield  $\hat{S}$ , which dynamically generates an effective  $\mu$ -term. This predicts an SM-like 125 GeV Higgs under all constraints and with low fine-tuning [2, 3]. The fully constrained NMSSM (cNMSSM) contains none or only one more parameter than the CMSSM/mSUGRA.

Received 20 January 2020 Published online 27 April 2020

\* Supported by the National Natural Science Foundation of China (NNSFC) (11605123)

1) E-mail: zhujy@whu.edu.cn



Content from this work may be used under the terms of the Creative Commons Attribution 3.0 licence. Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI. Article funded by SCOAP<sup>3</sup> and published under licence by Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd

GRA, thus both of them are in tension with current experimental constraints including 125 GeV Higgs mass, high mass bound of gluino, muon  $g-2$ , and dark matter [4–23]. Hence, we consider the semi-constrained NMSSM (scNMSSM) that relaxes the unification of scalar masses by decoupling the square-masses of the Higgs bosons and the squarks/sleptons, which is also referred to as NMSSM with non-universal Higgs mass (NUHM) [24–26]. In the scNMSSM, the bino and wino are heavy because the high mass bound of gluino and the unification of gaugino masses at GUT scale, thus the light neutralinos and charginos can only be singlino-dominated or higgsino-dominated<sup>1)</sup>.

In recent years, ATLAS and CMS collaborations have carried out numerous searches for SUSY particles, which pushed the gluino and squarks masses bounds in simple models up to several hundred GeV and even the TeV scale. Meanwhile, it remains possible for the electroweakino sector to be very light. The electroweakino sector of NMSSM was studied in [27–44], among which different search channels were provided, such as multi-leptons [35, 36] and jets with missing transverse momentum ( $\vec{p}_T$ ) [37]. These motivated us to verify the current status of higgsino in special SUSY models such as the scNMSSM, under direct-search constraints and their possibility of discovery by detailed simulation.

In this study, we discuss the light higgsino-dominated NLSPs (next-to-lightest supersymmetric particles) in the scNMSSM. We use the scenario of singlino-dominated  $\tilde{\chi}_1^0$  and SM-like  $h_2$  in the scan result in our former work on scNMSSM [26], where we considered the constraints including theoretical constraints of vacuum stability and Landau pole, experimental constraints of Higgs data, muon  $g-2$ ,  $B$  physics, dark matter relic density, and direct searches, etc. Thus, in this scenario the  $\tilde{\chi}_1^+$  and  $\tilde{\chi}_{2,3}^0$  are higgsino-dominated, light and mass-degenerated NLSPs. We first investigate the constraints on these NLSPs, including searching for SUSY particles at the LHC Run-I and Run-II. We employ Monte Carlo algorithm to perform detailed simulations to impose these constraints. Then, we discuss the possibility of discovering the higgsino-dominated NLSPs in the future at the high-luminosity LHC (HL-LHC).

This paper is organized as follows. First, in Section 2, we briefly introduce the model of NMSSM and scNMSSM, especially in the Higgs and electroweakino sector. Later in Section 3, we discuss the constraints to the light higgsino-dominated NLSPs, and the possibility of discovering them at the HL-LHC. Finally, we draw our conclusions in Section 4.

## 2 Introduction to NMSSM and scNMSSM

The superpotential of NMSSM with  $\mathbb{Z}_3$  symmetry :

$$W_{\text{NMSSM}} = W_{\text{MSSM}}|_{\mu=0} + \lambda \hat{S} \hat{H}_u \cdot \hat{H}_d + \frac{\kappa}{3} \hat{S}^3, \quad (1)$$

where superfields  $\hat{H}_u$  and  $\hat{H}_d$  are two complex doublet superfields, superfield  $\hat{S}$  is the singlet superfield, coupling constants  $\lambda$  and  $\kappa$  are dimensionless, and  $W_{\text{MSSM}}|_{\mu=0}$  presents the Yukawa couplings of the  $\hat{H}_u$  and  $\hat{H}_d$  to the quark and lepton superfields. In electroweak symmetry breaking, the scalar component of superfields  $\hat{H}_u$ ,  $\hat{H}_d$ , and  $\hat{S}$  obtain their vacuum expectation values (VEVs)  $v_u$ ,  $v_d$ , and  $v_s$  respectively. The relations between the VEVs are

$$\tan\beta = v_u/v_d, \quad v = \sqrt{v_u^2 + v_d^2} = 174 \text{ GeV}, \quad \mu_{\text{eff}} = \lambda v_s, \quad (2)$$

where the  $\mu_{\text{eff}}$  is the mass scale of a higgsino, like in the MSSM. In the following, for the sake of convenience, we refer to  $\mu_{\text{eff}}$  as  $\mu$ .

The soft SUSY breaking terms in the NMSSM are only different from the MSSM in several terms:

$$-\mathcal{L}_{\text{NMSSM}}^{\text{soft}} = -\mathcal{L}_{\text{MSSM}}^{\text{soft}}|_{\mu=0} + m_S^2 |S|^2 + \lambda A_\lambda S H_u \cdot H_d + \frac{1}{3} \kappa A_\kappa S^3 + \text{h.c.}, \quad (3)$$

where  $S$ ,  $H_u$ , and  $H_d$  are the scalar components of the superfields, the  $m_S^2$  is the soft SUSY breaking mass for singlet field  $S$ , and the trilinear coupling constants  $A_\lambda$  and  $A_\kappa$  have a mass dimension.

In the semi-constrained NMSSM (scNMSSM), the Higgs sector are considered non-universal, that is, the Higgs soft mass  $m_{H_u}^2$ ,  $m_{H_d}^2$ , and  $m_S^2$  are allowed to be different from  $M_0^2 + \mu^2$ , and the trilinear couplings  $A_\lambda$ ,  $A_\kappa$  can be different from  $A_0$ . Hence, in the scNMSSM, the complete parameter sector is usually chosen as:

$$\lambda, \kappa, \tan\beta, \mu, A_\lambda, A_\kappa, A_0, M_{1/2}, M_0. \quad (4)$$

### 2.1 Higgs sector of NMSSM and scNMSSM

When the electroweak symmetry breaks, the scalar component of superfields  $\hat{H}_u$ ,  $\hat{H}_d$ , and  $\hat{S}$  can be written as

$$\begin{aligned} H_u &= \begin{pmatrix} H_u^+ \\ v_u + \frac{H_u^R + iH_u^I}{\sqrt{2}} \end{pmatrix}, \\ H_d &= \begin{pmatrix} v_d + \frac{H_d^R + iH_d^I}{\sqrt{2}} \\ H_d^- \end{pmatrix}, \\ S &= v_s + \frac{S^R + iS^I}{\sqrt{2}}, \end{aligned} \quad (5)$$

1) The behaviors of the cNMSSM under recent constraints is similar as the CMSSM, where the additional parameter  $\lambda$  is very small, and higgsino mass parameter  $\mu_{\text{eff}}$  is calculated to be very large, thus the higgsino-dominated and singlino-dominated neutralinos are very heavy.

where  $H_u^R$ ,  $H_d^R$ , and  $S^R$  are  $CP$ -even component fields,  $H_u^I$ ,  $H_d^I$ , and  $S^I$  are the  $CP$ -odd component fields, and the  $H_u^\pm$  and  $H_d^\pm$  are charged component fields

On the basis of  $(H_d^R, H_u^R, S^R)$ , the  $CP$ -even scalar mass matrix is [45]

$$\mathcal{L} \ni \frac{1}{2} (H_d^R, H_u^R, S^R) M_S^2 \begin{pmatrix} H_d^R \\ H_u^R \\ S^R \end{pmatrix}, \quad (6)$$

with

$$M_S^2 = \begin{pmatrix} M_A^2 s_\beta^2 + M_Z^2 c_\beta^2 & (2\lambda v^2 - M_A^2 - M_Z^2) s_\beta c_\beta & C c_\beta + C' s_\beta \\ (2\lambda v^2 - M_A^2 - M_Z^2) s_\beta c_\beta & M_A^2 c_\beta^2 + M_Z^2 s_\beta^2 & C s_\beta + C' c_\beta \\ C c_\beta + C' s_\beta & C s_\beta + C' c_\beta & M_{S, S^R S^R}^2 \end{pmatrix}, \quad (7)$$

where

$$M_A^2 = \frac{2\mu(A_\lambda + \kappa v_s)}{\sin 2\beta}, \quad (8)$$

$$C = 2\lambda^2 v v_s, \quad C' = \lambda v(A_\lambda - 2\kappa v_s), \quad (9)$$

$$M_{S, S^R S^R}^2 = \lambda A_\lambda \frac{v_u v_d}{v_s} + \kappa v_s(A_\lambda + 4\kappa v_s), \quad (10)$$

and  $s_\beta = \sin \beta$ ,  $c_\beta = \cos \beta$ . Indeed, here is a more common basis  $(H_1, H_2, S^R)$ , where

$$H_1 = H_u^R c_\beta - H_d^R s_\beta, \quad H_2 = H_u^R s_\beta + H_d^R c_\beta \quad (11)$$

and the  $H_2$  is the SM Higgs field. On the basis of  $(H_1, H_2, S^R)$ , the scalar mass matrix is different from Eq. (7). However, since the rotation of the basis does not touch the third component  $S^R$ , the  $M_{S, S^R S^R}^2$  will remain the same as in Eq. (10). The Higgs boson mass matrix  $M_{S'}^2$  on basis  $(H_1, H_2, S^R)$  is given by [46]

$$M_{S', H_1 H_1}^2 = M_A^2 + (m_Z^2 - \lambda^2 v^2) \sin^2 2\beta, \quad (12)$$

$$M_{S', H_1 H_2}^2 = -\frac{1}{2} (m_Z^2 - \lambda^2 v^2) \sin 4\beta, \quad (13)$$

$$M_{S', H_1 S^R}^2 = -\left( \frac{M_A^2}{2\mu / \sin 2\beta} + \kappa v_s \right) \lambda v \cos 2\beta, \quad (14)$$

$$M_{S', H_2 H_2}^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta, \quad (15)$$

$$M_{S', H_2 S^R}^2 = 2\lambda \mu v \left[ 1 - \left( \frac{M_A}{2\mu / \sin 2\beta} \right)^2 - \frac{\kappa}{2\lambda} \sin 2\beta \right], \quad (16)$$

$$M_{S', S^R S^R}^2 = \frac{1}{4} \lambda^2 v^2 \left( \frac{M_A}{\mu / \sin 2\beta} \right)^2 + \kappa v_s A_\lambda + 4(\kappa v_s)^2 - \frac{1}{2} \lambda \kappa v^2 \sin 2\beta. \quad (17)$$

And comparing Eq. (17) with Eq. (10), it's not hard to get  $M_{S', S^R S^R}^2 = M_{S, S^R S^R}^2$ .

To obtain the physical  $CP$ -odd scalar Higgs bosons, the Higgs fields may be rotated,

$$A = H_u^I c_\beta + H_d^I s_\beta. \quad (18)$$

Subsequently, the Goldstone mode is dropped off, and the  $CP$ -odd scalar mass matrix on the basis of  $(A, S^I)$  becomes [45]

$$\mathcal{L} \ni \frac{1}{2} (A, S^I) M_P^2 \begin{pmatrix} A \\ S^I \end{pmatrix} \quad (19)$$

with

$$M_P^2 = \begin{pmatrix} M_A^2 & \lambda v(A_\lambda - 2\kappa v_s) \\ \lambda v(A_\lambda - 2\kappa v_s) & M_{P, S^I S^I}^2 \end{pmatrix}, \quad (20)$$

where

$$M_{P, S^I S^I}^2 = \lambda(A_\lambda + 4\kappa v_s) \frac{v_u v_d}{v_s} - 3\kappa v_s A_\lambda. \quad (21)$$

The mass eigenstates of the  $CP$ -even Higgs  $h_i$  ( $i = 1, 2, 3$ ) and the  $CP$ -odd Higgs  $A_i$  ( $i = 1, 2$ ) are obtained by

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = S_{ij} \begin{pmatrix} H_1 \\ H_2 \\ S^R \end{pmatrix}, \quad \begin{pmatrix} a_1 \\ a_2 \end{pmatrix} = P_{ij} \begin{pmatrix} A \\ S^I \end{pmatrix}, \quad (22)$$

where the matrix  $S_{ij}$  can diagonalize the mass matrix  $M_{S'}^2$ , and the matrix  $P_{ij}$  can diagonalize the mass matrix  $M_P^2$ .

## 2.2 Electroweakino sector of NMSSM and scNMSSM

In the NMSSM, there are five neutralinos  $\tilde{\chi}_i^0$  ( $i = 1, 2, 3, 4, 5$ ), which are a mixture of  $\tilde{B}$  (bino),  $\tilde{W}^3$  (wino),  $\tilde{H}_d^0$ ,  $\tilde{H}_u^0$  (higgsinos), and  $\tilde{S}$  (singlino). On the gauge-eigenstate basis  $\psi^0 = (\tilde{B}, \tilde{W}^3, \tilde{H}_d^0, \tilde{H}_u^0, \tilde{S})$ , the neutralino mass matrix takes the form [45]

$$M_{\tilde{\chi}^0} = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z & 0 \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z & 0 \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu & -\lambda v_d \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 & -\lambda v_u \\ 0 & 0 & -\lambda v_d & -\lambda v_u & 2\kappa v_s \end{pmatrix}, \quad (23)$$

where  $s_\beta = \sin\beta$ ,  $c_\beta = \cos\beta$ ,  $s_W = \sin\theta_W$ ,  $c_W = \cos\theta_W$ . To obtain the mass eigenstates, one may diagonalize the neutralino mass matrix  $M_{\tilde{\chi}^0}$

$$N^* M_{\tilde{\chi}^0} N^{-1} = M_{\tilde{\chi}^0}^D = \text{Diag}(m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_3^0}, m_{\tilde{\chi}_4^0}, m_{\tilde{\chi}_5^0}), \quad (24)$$

where  $M_{\tilde{\chi}^0}^D$  is the diagonal mass matrix, and the order of eigenvalues is  $m_{\tilde{\chi}_1^0} < m_{\tilde{\chi}_2^0} < m_{\tilde{\chi}_3^0} < m_{\tilde{\chi}_4^0} < m_{\tilde{\chi}_5^0}$ . Meanwhile, the mass eigenstates are obtained

$$\begin{pmatrix} \tilde{\chi}_1^0 \\ \tilde{\chi}_2^0 \\ \tilde{\chi}_3^0 \\ \tilde{\chi}_4^0 \\ \tilde{\chi}_5^0 \end{pmatrix} = N_{ij} \begin{pmatrix} \tilde{B} \\ \tilde{W}^0 \\ \tilde{H}_d \\ \tilde{H}_u \\ \tilde{S} \end{pmatrix}. \quad (25)$$

In the scNMSSM, bino and wino were constrained to be very heavy because of the high mass bounds of gluino and the universal gaugino mass at the GUT scale. Thus, they can be decoupled from the light sector. Then, the following relations for the  $N_{ij}$  are found [47, 48]:

$$N_{i3} : N_{i4} : N_{i5} = \left[ \frac{m_{\tilde{\chi}_i^0}}{\mu} s_\beta - c_\beta \right] : \left[ \frac{m_{\tilde{\chi}_i^0}}{\mu} c_\beta - s_\beta \right] : \frac{\mu - m_{\tilde{\chi}_i^0}}{\lambda v}. \quad (26)$$

We assume that the lightest neutralino  $\tilde{\chi}_1^0$  is the lightest supersymmetric particle (LSP) and makes up of the cosmic dark matter. If the LSP  $\tilde{\chi}_1^0$  satisfies  $N_{15}^2 > 0.5$ , we call it singlino-dominated. The coupling of such an LSP with the  $CP$ -even Higgs bosons is given by [47, 48]

$$C_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} = \sqrt{2}\lambda[S_{i1}N_{15}(N_{13}c_\beta - N_{14}s_\beta) + S_{i2}N_{15} \times (N_{14}c_\beta + N_{13}s_\beta) + S_{i3}(N_{13}N_{14} - \frac{\kappa}{\lambda}N_{15}^2)]. \quad (27)$$

In the singlino-dominated-LSP scenario, assuming  $N_{11} = N_{12} = 0$ , the mass of LSP can be written as  $m_{\tilde{\chi}_1^0} \approx M_{\tilde{\chi}^0, \tilde{S}\tilde{S}} = 2\kappa v_s$ . From Eq. (23), Eq. (10), and Eq. (21), one can find the sum rule [49, 50]:

$$M_{\tilde{\chi}^0, \tilde{S}\tilde{S}}^2 = 4\kappa^2 v_s^2 = M_{\tilde{S}, \tilde{S}^R \tilde{S}^R}^2 + \frac{1}{3}M_{\tilde{P}, \tilde{S}' \tilde{S}'}^2 - \frac{4}{3}v_u v_d \left( \frac{\lambda^2 A_\lambda}{\mu} + \kappa \right). \quad (28)$$

In the case where  $h_1$  singlet-like,  $\tan\beta$  sizable,  $\lambda, \kappa$ , and  $A_\lambda$  are not too large, this equation can become

$$m_{\tilde{\chi}_1^0}^2 \approx m_{h_1}^2 + \frac{1}{3}m_{a_1}^2. \quad (29)$$

The chargino sector in the NMSSM is similar to the neutralino sector. The charged higgsino  $\tilde{H}_u^+$ ,  $\tilde{H}_d^-$  (with mass scale  $\mu$ ) and the charged gaugino  $\tilde{W}^\pm$  (with mass scale  $M_2$ ) can combine, forming two couples of physical chargino  $\chi_1^\pm, \chi_2^\pm$ . On the gauge-eigenstate basis  $(\tilde{W}^+, \tilde{H}_u^+, \tilde{W}^-, \tilde{H}_d^-)$ , the chargino mass matrix is given by [45]

$$M_C = \begin{pmatrix} 0 & X^T \\ X & 0 \end{pmatrix}, \quad \text{where } X = \begin{pmatrix} M_2 & \sqrt{2}s_\beta m_W \\ \sqrt{2}c_\beta m_W & \mu \end{pmatrix}. \quad (30)$$

To obtain the chargino mass eigenstates, one can use two unitary matrices to diagonalize the chargino mass matrix by

$$U^* X V^{-1} = M_{\tilde{\chi}^\pm}^D = \text{Diag}(m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_2^\pm}), \quad (31)$$

where  $M_{\tilde{\chi}^\pm}^D$  indicates the diagonal mass matrix, and the order of eigenvalues is  $m_{\tilde{\chi}_1^\pm} < m_{\tilde{\chi}_2^\pm}$ . Meanwhile, we obtain the mass eigenstates

$$\begin{pmatrix} \tilde{\chi}_1^\pm \\ \tilde{\chi}_2^\pm \end{pmatrix} = V_{ij} \begin{pmatrix} \tilde{W}^\pm \\ \tilde{H}_u^\pm \end{pmatrix}, \quad \begin{pmatrix} \tilde{\chi}_1^\mp \\ \tilde{\chi}_2^\mp \end{pmatrix} = U_{ij} \begin{pmatrix} \tilde{W}^\mp \\ \tilde{H}_d^\mp \end{pmatrix}. \quad (32)$$

In the scNMSSM, since  $M_2 \gg \mu$ ,  $\chi_1^\pm$  can be higgsino-dominated, with a mass of approximately  $\mu$ . With  $\chi_1^0$  is singlino-dominated,  $\chi_{2,3}^0$  can be higgsino-dominated, with masses that are nearly degenerate and approximately  $\mu$  and  $N_{23}^2 + N_{24}^2 > 0.5$ . When  $\mu$  is not large, i.e., smaller than the mass of other particles, the nearly-degenerate  $\chi_1^\pm$  and  $\chi_{2,3}^0$  are referred to as the next-to-lightest SUSY particles (NLSPs). In this study, we focus on the detection of the higgsino-dominated NLSPs ( $\chi_1^\pm$  and  $\chi_{2,3}^0$ ) in the scNMSSM.

### 3 Light Higgsino-dominated NLSPs in scNMSSM

We employ the scan result from our previous study on scNMSSM [26]; however, we only consider the surviving samples with singlino-dominated  $\chi_1^0$  ( $|N_{15}|^2 > 0.5$ ) as the LSP, and impose the SUSY search constraints with CheckMATE [51–53]. We perform the scan with the program NMSSMTools-5.4.1 [54–57], and considered the constraints there, including theoretical constraints of vacuum stability and the Landau pole, experimental constraints of Higgs data, muon g-2, B physics, dark matter relic density, and direct searches, etc.<sup>1)</sup> We also use HiggsBounds-5.1.1beta [58] to constrain the Higgs sector (with  $h_2$  as the SM-like Higgs and  $123 < m_{h_2} < 127$  GeV), and SModelS-v1.1.1 [59, 60] to constrain SUSY particles. The detail of the constraints can be found in Ref. [26]. The scanned spaces of the parameters are:

$$\begin{aligned} 0 < M_0 < 500 \text{ GeV}, \quad 0 < M_{1/2} < 2 \text{ TeV}, \quad |A_0| < 10 \text{ TeV}, \\ 100 < \mu < 200 \text{ GeV}, \quad 1 < \tan\beta < 30, \quad 0.3 < \lambda < 0.7, \\ 0 < \kappa < 0.7, \quad |A_\lambda| < 10 \text{ TeV}, \quad |A_\kappa| < 10 \text{ TeV}. \end{aligned} \quad (33)$$

As shown in Ref. [26], in the surviving parameter space,

- The gluino is heavier than 1.5 TeV, i.e.,  $M_{1/2}$  at

1) For the dark matter relic density, we only consider the upper bound, that is  $\Omega h^2 \leq 0.131$ , considering there may be also other sources of dark matter.



GUT scale, or  $M_3/2.4 \simeq M_2/0.8 \simeq M_1/0.4$  at  $M_{\text{SUSY}}$  scale due to RGE runnings, and it is larger than approximately 700 GeV.

- Because of RGE runnings, including  $M_3$ , the squarks can be heavy. Even the lightest squarks, e.g.,  $\tilde{t}_1$ , are heavier than approximately 500 GeV.

- Because of RGE runnings, including  $M_2$  and  $M_1$ , the sleptons of the first two generations and all neutrinos are heavier than approximately 300 GeV, only  $\tilde{\tau}_1$  can be lighter, at approximately 100 GeV.

- The heavy charginos  $\tilde{\chi}_2^\pm$  are wino-like and heavier than approximately 560 GeV, the light charginos  $\tilde{\chi}_1^\pm$  are higgsino-like at 100 ~ 200 GeV.

- For the five neutralinos,  $\tilde{\chi}_5^0$  is wino-like and heavier than 560 GeV, the bino-dominated neutralino is heavier than 280 GeV, the higgsino-dominated neutralinos are 100 ~ 200 GeV, while the singlino-dominated neutralino can be 60 ~ 400 GeV.

- For the Higgs sector,  $h_2$  represents the SM-like Higgs at 123 ~ 127 GeV,  $h_1$  is singlet-dominated and lighter than 123 GeV, the light  $CP$ -odd Higgs is singlet-dominated, but it can be lighter or heavier than 125 GeV. Since  $h_2$  is SM-like, the Higgs invisible decay caused by  $m_{\tilde{\chi}_1^0} \simeq 60$  GeV is at most about 20%, so as the Higgs exotic decays caused by  $m_{h_{1,a}} \lesssim 60$  GeV.

In this study, we choose the surviving samples with  $\tilde{\chi}_1^0$  LSP as singlino-dominated, and higgsino-dominated neutralino and chargino ( $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_{2,3}^0$ ) as NLSPs. Thus, the samples with higgsino-dominated neutralinos as LSPs, or  $\tilde{\tau}_1$  as NLSP, are discarded. In the following, we focus on the higgsino-dominated NLSPs in the scNMSSM, considering its constraints from direct search results at the LHC Run I and Run II, its production and decay, and its possibility of discovery at the HL-LHC in the future.

### 3.1 Constraints from direct SUSY searches with CheckMATE

Furthermore, we use CheckMATE 2.0.26 [51–53] to impose these constraints of direct SUSY search results at the LHC, using all data at Run I and up to 36 fb<sup>-1</sup> data at Run II [61–66]. For masses of 100 ~ 200 GeV, the cross-sections of the higgsino-dominated NLSPs can be sizeable, thus we pay special attention to the NLSPs.

First, We use MadGraph5\_aMC@NLO 2.6.6 [67] to generate three types of tree level processes at 8 TeV and 13 TeV:

$$pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-, \quad pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_{2,3}^0, \quad pp \rightarrow \tilde{\chi}_{2,3}^0 \tilde{\chi}_{2,3}^0. \quad (34)$$

Since the cross-sections by the MadGraph are at the tree

level, we multiply them by a NLO K-factor calculated with the Prospino2 [68]. Then, we use PYTHIA 8.2 [69] to deal with particle decay, parton showering, and hadronization, and we use Delphes 3.4.1 [70] to simulate the detector response. The anti- $k_T$  algorithm [71] is employed for jet clustering.

After the simulation, we obtain a ‘.root’ file. We use the CheckMATE2 to read this ‘.root’ file. Then, we apply the same cuts in signal regions of the CMS and ATLAS experiments at 8 TeV and 13 TeV, using analysis cards that have been implemented in CheckMATE2. At the last step, with the CheckMATE2, we obtain an  $r$ -value for each sample, which is defined as

$$r \equiv \frac{S - 1.64\Delta S}{S_{\text{Exp}}^{95}}, \quad (35)$$

where  $S$  is the total number of expected signal events,  $\Delta S$  is the uncertainty of  $S$ , and  $S_{\text{Exp}}^{95}$  is the experimentally measured at the 95% confidence limit of the signal events number. Hence, a model can be considered excluded at the 95% confidence level, if  $r \geq 1$ . If the  $r \geq 1$  in only one signal region, the model can also be excluded. We obtain  $r_{\text{max}}$ , the maximal value of  $r$  in different signal regions. The model is excluded if  $r_{\text{max}} \geq 1$ .

After using CheckMATE to verify our surviving samples, we note that most of the samples excluded are by the CMS analysis in multilepton final states at 13 TeV LHC with 35.9 fb<sup>-1</sup> data [61–64, 72]<sup>1)</sup>. We confirmed that the relevant mechanism is  $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$  produced and each decaying to two bodies. Since the sleptons are heavier, the  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  mainly decay to  $\tilde{\chi}_1^0$  LSP plus a  $W$ , or  $Z$ , or Higgs boson. The most effective processes excluding the samples are  $pp \rightarrow \tilde{\chi}_1^\pm (W^\pm \tilde{\chi}_1^0) \tilde{\chi}_2^0 (Z \tilde{\chi}_1^0)$  and  $pp \rightarrow \tilde{\chi}_1^\pm (W^\pm \tilde{\chi}_1^0) \tilde{\chi}_2^0 (h_2 \tilde{\chi}_1^0)$ .

The search strategy for these two processes is three or more leptons and a large  $\vec{p}_T$  in the final state, since these channels are relatively cleaner than the jet channels at the LHC. The CMS searches related to our processes include the following signal regions (SR) SR-A, SR-C, and SR-F

- **SR-A:** events with three light leptons ( $e$  or  $\mu$ ), two of which form an opposite sign same-flavor (OSSF) pair. The SR-A is divided into 44 bins, according to the invariant mass of an OSSF pair  $M_{\ell\ell}$ , third lepton's transverse mass  $M_T$ , and missing transverse momentum  $\vec{p}_T$ . The transverse mass  $M_T$  is defined as

$$M_T = \sqrt{2p_T^\ell p_T [1 - \cos(\Delta\phi)]}, \quad (36)$$

where  $\Delta\phi$  is the angle between  $\vec{p}_T^\ell$  and  $\vec{p}_T$ .

1) The CMS collaboration has not released update results with more data in the multilepton channel up to now. The cut scheme of ATLAS analysis in this channel is different, and we checked that the ATLAS result with 13.3 fb<sup>-1</sup> data implemented in the CheckMATE is much weaker than the CMS result in constraining our samples, and thus even with 139 fb<sup>-1</sup> data it has no significant impact on our final conclusion. We also have considered the CMS analysis for the compressed spectrum with 12.9 fb<sup>-1</sup> data, and find that even with the current 139 fb<sup>-1</sup> data it can not constrain our samples more.

• **SR-C:** events with two light leptons ( $e$  or  $\mu$ ) forming an OSSF pair, and one  $\tau_h$  candidate. The SR-C is divided into 18 bins, according to the invariant mass  $M_{\ell\ell}$ , the two-lepton ‘transverse mass’  $M_{T2}(\ell_1, \ell_2)$  [73–75] instead of  $M_T$  on the off- $Z$  regions, and the  $\vec{p}_T$ . The  $M_{T2}$  is defined as

$$M_{T2} = \min_{\vec{p}_T^1 + \vec{p}_T^2 = \vec{p}_T} \left[ \max \left\{ M_T(\vec{p}_T^1, \vec{p}_T^1), M_T(\vec{p}_T^2, \vec{p}_T^2) \right\} \right], \quad (37)$$

where  $\vec{p}_T^1$  and  $\vec{p}_T^2$  are the transverse momenta for the two leptons, respectively, while  $\vec{p}_T^1$  and  $\vec{p}_T^2$  depict the random two components of the missing transverse momenta  $\vec{p}_T$ . This is used to suppress the SM background, since the large  $t\bar{t}$  background is at low  $M_{T2}$ .

• **SR-F:** events with one light lepton ( $e$  or  $\mu$ ) plus two  $\tau_h$  candidates. SR-F is divided into 12 bins, according to  $M_{\ell\ell}$ ,  $M_{T2}(\ell, \tau_1)$ , and the  $\vec{p}_T$ .

Recently after 2017, ATLAS and CMS collaborations have released several new search results with more Run-II data up to 139 fb $^{-1}$  in channels as  $2\ell + \vec{p}_T$  [76–81],  $3\ell + \vec{p}_T$  [72] (see also non-SUSY interpretations of multilepton anomalies in Refs. [82–84]),  $2\gamma + \vec{p}_T$  [85, 86], Higgs +  $\vec{p}_T$  [87], and Higgs +  $\ell + \vec{p}_T$  [88], and for compressed mass spectrum [89, 90]. In their analyses, they considered simple models, where purely higgsino or wino NLSP produced in pair, each decaying to  $\tilde{\chi}_1^0$  plus  $h$ ,  $Z$ , or  $W^\pm$  in 100%, such that the results do not apply to our samples directly. Imposing these new constraints is complex, and we plan to address them in our future studies.

### 3.2 Production and decay of Higgsino-dominated NLSPs

For the surviving samples, we first investigated the production cross-sections of  $\tilde{\chi}_1^+ \tilde{\chi}_2^0$ ,  $\tilde{\chi}_1^+ \tilde{\chi}_3^0$ ,  $\tilde{\chi}_1^- \tilde{\chi}_2^0$ ,  $\tilde{\chi}_1^- \tilde{\chi}_3^0$ ,  $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ ,  $\tilde{\chi}_2^0 \tilde{\chi}_3^0$ ,  $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ , and  $\tilde{\chi}_3^0 \tilde{\chi}_3^0$  at the 14 TeV LHC. Since the NLSPs are higgsino-dominated, the cross-sections are not significantly different from those of the pure higgsino production. Here, we only revise some of the old conclusions.

• The cross-sections decrease quickly when masses increase, for the partonic Mandelstam variable  $\hat{s}$  increase, and the parton fluxes decrease.

• The cross-section of  $\tilde{\chi}_1^+ \tilde{\chi}_i^0$  is approximately two times  $\tilde{\chi}_1^- \tilde{\chi}_i^0$ , for both  $i = 2$  and  $3$ . The reason is that the LHC is proton-proton collider, and the parton distribution functions (PDF) for up quark are larger than for the down quark.

• The cross-sections of  $\tilde{\chi}_2^0 \tilde{\chi}_2^0$  and  $\tilde{\chi}_3^0 \tilde{\chi}_3^0$  are very small, in the range of a few fb. The reason is that the squarks are very heavy, such that  $pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_i^0$  are mainly produced from the  $s$  channel through  $Z$  boson resonance. The coupling of  $Z - \tilde{\chi}_i^0 - \tilde{\chi}_j^0$  is given by

$$C_{Z\tilde{\chi}_i^0\tilde{\chi}_j^0} = -\frac{i}{2}(g_1 s_W + g_2 c_W)(N_{j3}^* N_{i3} - N_{j4}^* N_{i4})(\gamma_\mu P_L) + \frac{i}{2}(g_1 s_W + g_2 c_W)(N_{i3}^* N_{j3} - N_{i4}^* N_{j4})(\gamma_\mu P_R), \quad (38)$$

where the matrix  $N$  is the neutralino mixing matrix. Since  $\tilde{\chi}_{2,3}^0$  are higgsino-dominated,  $m_{\tilde{\chi}_{2,3}^0} \simeq \mu$ , from Eq. (30) we have  $N_{i3} : N_{i4} \approx -1$ , such that  $(|N_{i3}|^2 - |N_{i4}|^2) \approx 0$ , and  $\sigma(pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_i^0) \approx 0$ .

The decay branching ratios of the NLSPs are shown in Fig. 1 and Fig. 2. From the left plots in Fig. 1, we see that the chargino  $\tilde{\chi}_1^\pm$  decays to  $\tilde{\chi}_1^0$  plus a  $W$  boson in 100%: when the mass difference between  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_1^0$  is larger than  $m_W$ , the  $W$  boson is a real one; whereas when the mass difference is insufficient, the  $W$  boson is a virtual one, or the decay is a three-body decay. The low mass difference is negative for us to search for the SUSY particles, since the leptons coming from a virtual  $W$  boson are very soft and hard to detect.

The main decay modes of the neutralino  $\tilde{\chi}_i^0$  ( $i = 2, 3$ ) are to a  $\tilde{\chi}_1^0$  plus a  $Z$  boson or a Higgs boson. In the middle and right plots of Fig. 1 and in Fig. 2, we show the branching ratios of the neutralinos  $\tilde{\chi}_i^0$  on the plane of  $m_{\tilde{\chi}_i^0}$  vs  $m_{\tilde{\chi}_1^0}$ , where  $i = 2, 3$ . In these plots, we use the dashed line,  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} = m_Z$ , and the dotted line,  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} = m_{h_2}$ , dividing the plane into three parts.

• **Case I:** In the region  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} < m_Z$ , the neutralino  $\tilde{\chi}_i^0$  can only decay to  $\tilde{\chi}_1^0$  plus a virtual  $Z$  boson or a light Higgs boson  $h_1$ . We observe from the lower middle plot of Fig. 1 and the upper left plot of Fig. 2, the  $\tilde{\chi}_2^0$  mainly decays to a virtual  $Z$  boson plus a  $\tilde{\chi}_1^0$ , with only a small fraction to the light Higgs boson  $h_1$  plus  $\tilde{\chi}_1^0$ . In contrast, the lower left plot of Fig. 2 shows that the  $\tilde{\chi}_3^0$  mainly decays to the light Higgs boson  $h_1$ .

• **Case II:** In the region  $m_Z \leq m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} < m_{h_2}$ , the neutralino  $\tilde{\chi}_i^0$  can decay to  $\tilde{\chi}_1^0$  plus a real  $Z$  boson or a light Higgs boson  $h_1/a_1$ . As shown in the upper middle plot of Fig. 1, the  $\tilde{\chi}_2^0$  mainly decay to a real  $Z$  boson plus  $\tilde{\chi}_1^0$ . According to the upper right plot of Fig. 1 and lower left plot of Fig. 2, the  $\tilde{\chi}_3^0$  mainly decays to a light Higgs boson  $h_1$  plus  $\tilde{\chi}_1^0$ .

• **Case III:** In the region  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} \geq m_{h_2}$ , all these decay channels are opened. The  $\tilde{\chi}_2^0$  mainly decays to  $\tilde{\chi}_1^0$  and a 125 GeV SM-like Higgs boson  $h_2$ , while  $\tilde{\chi}_3^0$  mainly decays to  $\tilde{\chi}_1^0$  and  $Z$  bosons.

In the channel  $\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 Z$  ( $i = 2, 3$ ), like the  $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$ , when the mass difference is insufficient, the  $Z$  boson also becomes a virtual one. In the channel  $\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 H$  ( $i = 2, 3$ ), where the Higgs boson can be  $h_1$  or  $h_2$ , and  $h_2$  is the SM-like one. Both  $h_1$  and  $h_2$  mainly decay to  $b\bar{b}$ , thus the  $t\bar{t}$  background is sizable at the LHC. In the case that Higgs decay to  $WW$ ,  $ZZ$ , or  $\tau\tau$ , and  $W$  or  $Z$  decays leptonically, this might contribute to the mul-

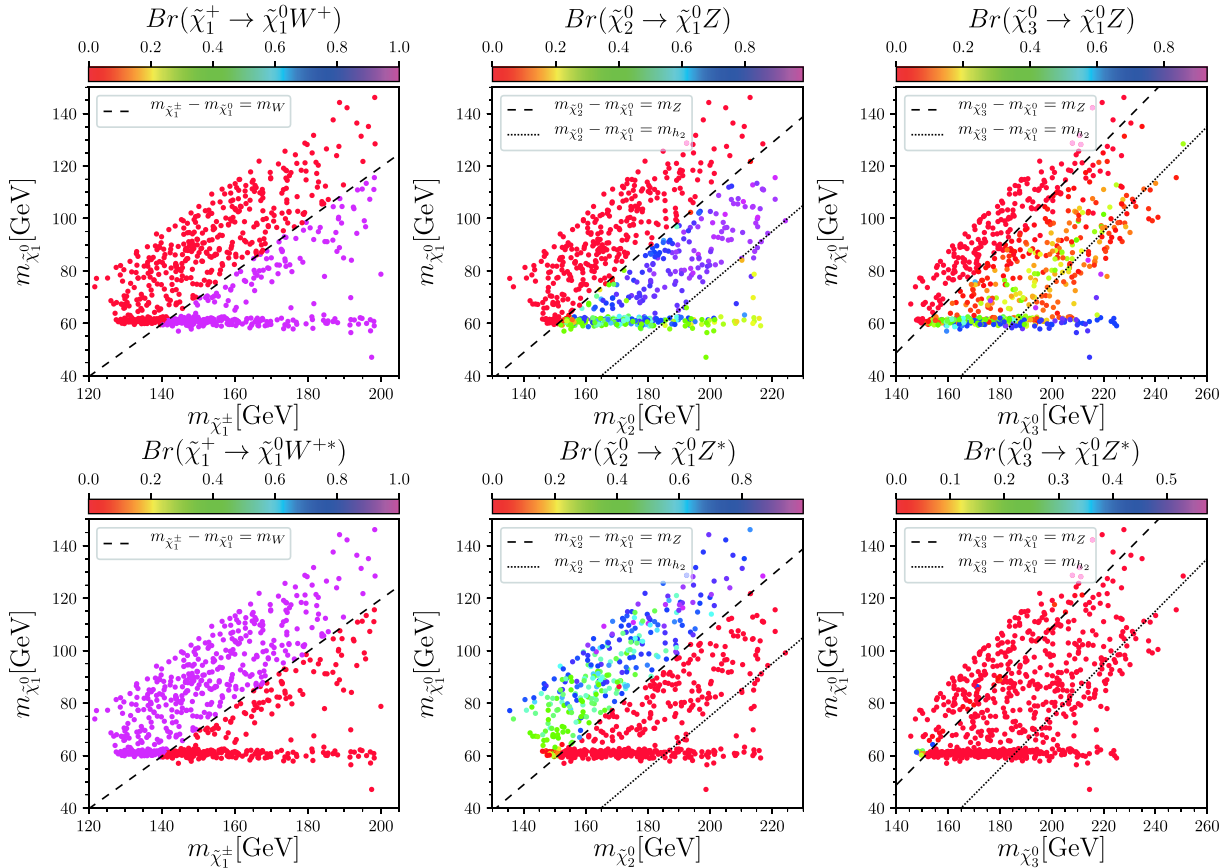


Fig. 1. (color online) Samples in  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_i^0}$  planes (left  $i = \pm$ , middle  $i = 2$ , and right  $i = 3$ ). Colors indicate branching ratios of chargino  $\tilde{\chi}_1^\pm$  to  $\tilde{\chi}_1^0$  plus  $W$  boson, and neutralino  $\tilde{\chi}_{2,3}^0$  to  $\tilde{\chi}_1^0$  plus  $Z$  boson, respectively. In the upper panel, the  $W/Z$  boson is a real one, and the decay is real two-body decay; while in the lower panel the  $W/Z$  boson is a virtual one and the decay is virtual three-body decay.

tilepton final state. Since the light Higgs  $h_1$  is highly singlet-dominated, the  $\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 h_1$  is very hard to contribute to the multilepton signal regions. Thus only the  $\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 h_2$  can visibly contribute to the multilepton signal regions.

Notably, when heavier neutralinos decay to the  $\tilde{\chi}_1^0$  LSP, the  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_3^0$  behave differently. Especially in the case II,  $\tilde{\chi}_2^0$  prefers to decay to a  $Z$  boson and  $\tilde{\chi}_1^0$ ,  $\text{Br}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 Z) > \text{Br}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 h_1)$ ; while  $\tilde{\chi}_3^0$  tends to decay to a light Higgs boson  $h_1$  and  $\tilde{\chi}_1^0$ ,  $\text{Br}(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 Z) < \text{Br}(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 h_1)$ . The couplings  $C_{h_1 \tilde{\chi}_2^0 \tilde{\chi}_1^0}$  and  $C_{h_1 \tilde{\chi}_3^0 \tilde{\chi}_1^0}$  can be written as

$$C_{h_1 \tilde{\chi}_2^0 \tilde{\chi}_1^0} \sim \frac{\lambda(N_{14}N_{23} + N_{13}N_{24})S_{13}}{\sqrt{2}} - \sqrt{2}\kappa N_{15}N_{25}S_{13}, \quad (39)$$

$$C_{h_1 \tilde{\chi}_3^0 \tilde{\chi}_1^0} \sim \frac{\lambda(N_{14}N_{33} + N_{13}N_{34})S_{13}}{\sqrt{2}} - \sqrt{2}\kappa N_{15}N_{35}S_{13}, \quad (40)$$

where the  $N_{11}$ ,  $N_{12}$ ,  $N_{21}$ ,  $N_{22}$ ,  $N_{31}$ , and  $N_{32}$  was set to 0 since the wino and bino are very heavy and decoupled in the scNMSSM, and the  $S_{11}$  and  $S_{12}$  was set to 0 since  $|S_{13}| \gg |S_{11}|, |S_{12}|$ .  $\lambda/\sqrt{2} \ll 1$  and  $\sqrt{2}\kappa \ll 1$ , such that the couplings  $C_{h_1 \tilde{\chi}_2^0 \tilde{\chi}_1^0}$  and  $C_{h_1 \tilde{\chi}_3^0 \tilde{\chi}_1^0}$  are both very small and roughly the same. Meanwhile, the couplings  $C_{Z \tilde{\chi}_2^0 \tilde{\chi}_1^0}$  and

$C_{Z \tilde{\chi}_3^0 \tilde{\chi}_1^0}$  can be different from each other according to Eq. (38), which can be approximated to

$$C_{Z \tilde{\chi}_2^0 \tilde{\chi}_1^0} \sim \frac{g_2}{c_W}(N_{13}N_{23} - N_{14}N_{24}), \quad (41)$$

$$C_{Z \tilde{\chi}_3^0 \tilde{\chi}_1^0} \sim \frac{g_2}{c_W}(N_{13}N_{33} - N_{14}N_{34}), \quad (42)$$

where the  $g_2/c_W \sim 1$ . When the two terms in Eq. (41) or Eq. (42) have different signs, and do not cancel each other, the couplings  $C_{Z \tilde{\chi}_i^0 \tilde{\chi}_1^0}$  can be considerably larger than  $C_{h_1 \tilde{\chi}_i^0 \tilde{\chi}_1^0}$ ; otherwise the cancel between the two terms can make  $C_{Z \tilde{\chi}_i^0 \tilde{\chi}_1^0}$  smaller than  $C_{h_1 \tilde{\chi}_i^0 \tilde{\chi}_1^0}$ . For some surviving samples,  $C_{Z \tilde{\chi}_3^0 \tilde{\chi}_1^0}$  have the cancellation between the two terms, and that leads to small  $\text{Br}(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 Z)$  and large  $\text{Br}(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 h_1)$ . Six benchmark points are listed in Table 1.

### 3.3 Possibility of future discovery at HL-LHC

We investigate the possibility of future detection of electroweakinos at the HL-LHC. We adopt the same analysis of the multilepton final state by CMS [61, 62], only increasing the integrated luminosity from  $35.9 \text{ fb}^{-1}$  to  $300 \text{ fb}^{-1}$ , to estimate the possibility of discovery in the future.

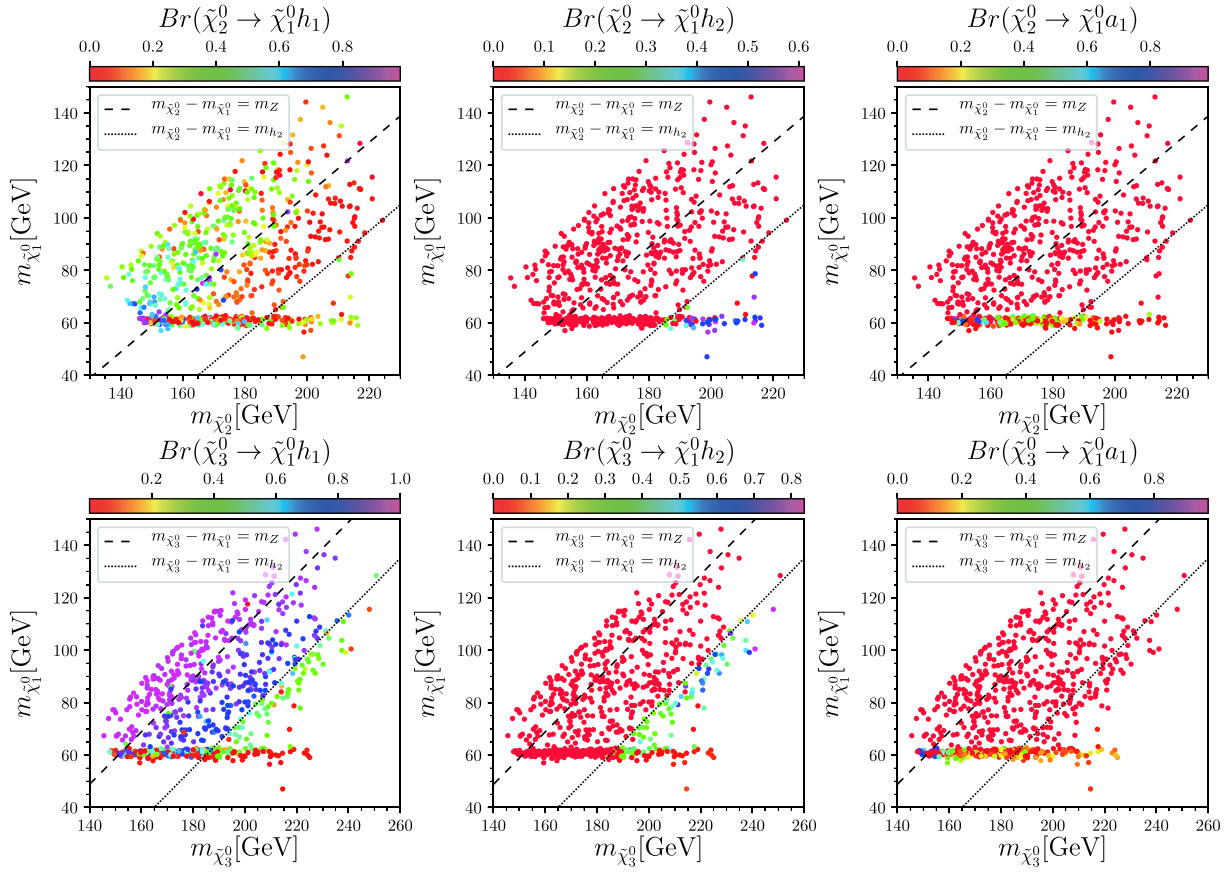


Fig. 2. (color online) Samples in  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_i^0}$  planes (upper  $i=2$ , lower  $i=3$ ). From left to the right, colors indicate the branching ratios  $Br(\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 h_1)$ ,  $Br(\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 h_2)$ , and  $Br(\tilde{\chi}_i^0 \rightarrow \tilde{\chi}_1^0 a_1)$ , respectively. The dashed and the dotted lines indicate that the mass differences,  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0}$ , is equal to  $m_Z$  and  $m_{h_2}$ , respectively.

Table 1. Masses and branching ratios for six benchmark points in scNMSSM. Signal significance in the last line is calculated with luminosity of  $300 \text{ fb}^{-1}$ , and with similar analysis of multi-lepton final state as in Refs. [61, 62].

	P1	P2	P3	P4	P5	P6
$m_{\tilde{\chi}_1^\pm} \text{ (GeV)}$	183	173	175	189	175	187
$m_{\tilde{\chi}_1^0} \text{ (GeV)}$	120	119	103	108	82	92
$m_{\tilde{\chi}_2^0} \text{ (GeV)}$	200	187	200	209	202	206
$m_{\tilde{\chi}_3^0} \text{ (GeV)}$	216	200	214	216	212	209
$Br(\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 W)$	0%	0%	0%	100%	100%	100%
$Br(\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 W^*)$	100%	100%	100%	0%	0%	0%
$Br(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 Z)$	0%	0%	90%	93%	94%	95%
$Br(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 Z^*)$	80%	65%	0%	0%	0%	0%
$Br(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 h_1)$	20%	35%	10%	7%	6%	5%
$Br(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 h_2)$	0%	0%	0%	0%	0%	0%
$Br(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 Z)$	1%	0%	1%	13%	13%	38%
$Br(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 Z^*)$	0%	0%	0%	0%	0%	0%
$Br(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 h_1)$	99%	100%	99%	87%	33%	62%
$Br(\tilde{\chi}_3^0 \rightarrow \tilde{\chi}_1^0 h_2)$	0%	0%	0%	0%	54%	0%
$ss = S / \sqrt{B} @ 300 \text{ fb}^{-1} \text{ (}\sigma\text{)}$	3.1	2.9	2.1	3.8	10.8	8.8



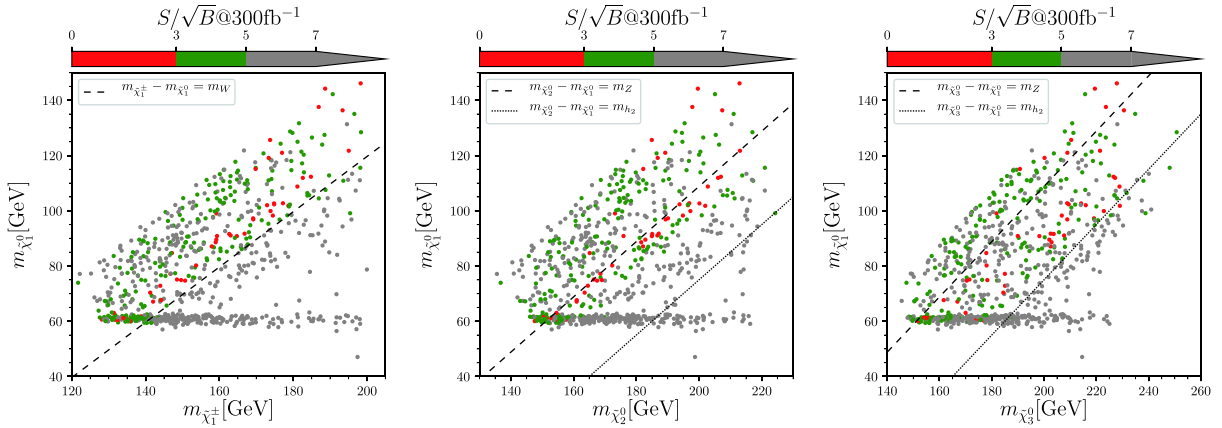


Fig. 3. (color online) Samples in  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_1^\pm}$  (left),  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_2^0}$  (middle),  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_3^0}$  (right) planes. Colors indicate the signal significance, where red represents  $ss < 3\sigma$ , green represents  $3\sigma < ss < 5\sigma$ , and gray represents  $ss > 5\sigma$ . In the left plane, the dashed line indicates the mass difference equal to  $m_W$ ,  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = m_W$ . In the middle and right planes, the dashed line and dotted line indicate the mass difference equal to  $m_Z$  and  $m_{h_2}$  respectively, that is,  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} = m_Z$  and  $m_{\tilde{\chi}_i^0} - m_{\tilde{\chi}_1^0} = m_{h_2}$ , where  $i = 2, 3$  for the middle and right planes, respectively.

We evaluate the signal significance by

$$ss = S / \sqrt{B}, \quad (43)$$

where  $S$  and  $B$  are the number of events from signal and background processes, respectively.

In Fig. 3, we show  $ss$  on the planes of  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_1^\pm}$ ,  $m_{\tilde{\chi}_2^0}$ , and  $m_{\tilde{\chi}_3^0}$  respectively. We can see that most of the samples can be evaluated at the  $5\sigma$  level when the mass difference between LSP  $\tilde{\chi}_1^0$  and NLSPs  $\tilde{\chi}_1^\pm, \tilde{\chi}_{2,3}^0$  is sufficient. However, there remain some samples that cannot be checked at level above 3 or 5 sigma. The main reason is that the mass spectra is compressed, such that the leptons from the decay of NLSPs are very soft. Because  $\vec{p}_T$  cut has to be very large at the LHC due to the large background, detecting soft particles is not easy. Combining with Figs. 1 and 2, we can learn the following facts:

- If  $\tilde{\chi}_1^\pm$  or  $\tilde{\chi}_i^0$  ( $i = 2, 3$ ) decays to a virtual vector boson, or in the area upon the dashed line in all the plots, the signal significance is less than  $5\sigma$ , and it is hard to check with  $300 \text{ fb}^{-1}$  data at the LHC.
- If  $\tilde{\chi}_1^\pm$  or  $\tilde{\chi}_2^0$  decays to a real vector boson, the area between the dashed and dotted line in left and middle plots, the signal significance can be larger than  $5\sigma$ , which is easy to verify with  $300 \text{ fb}^{-1}$  data at the LHC.
- The  $\tilde{\chi}_3^0$  decay to a light Higgs  $h_1$ , the area between the dashed and dotted line in the right plane, the signal significance is less than  $5\sigma$  for some samples. This is because the light Higgs  $h_1$  mainly decay to  $b\bar{b}$ , which is hard to distinguish from the background.
- If  $\tilde{\chi}_i^0$  ( $i = 2, 3$ ) decays to an SM-like Higgs, or in the area below the dotted line in middle and right plots, the signal significance is also larger than  $5\sigma$  for most samples.

For the samples with insufficient mass difference

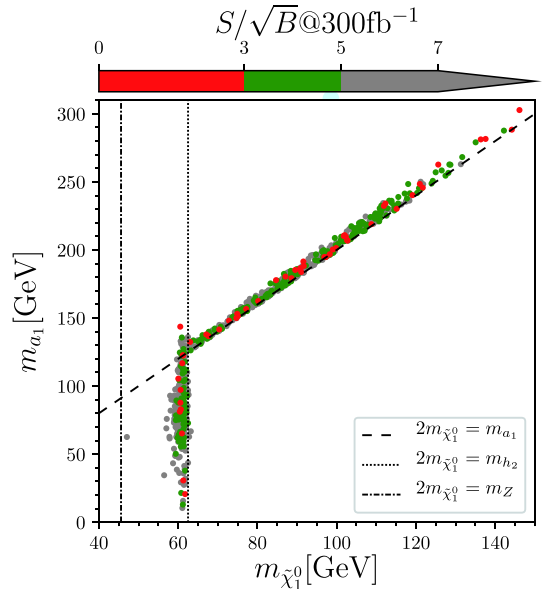


Fig. 4. (color online) Samples in  $m_{a_1}$  versus  $m_{\tilde{\chi}_1^0}$  plane. The color convention is the same as in Fig. 3. The dashed, dotted, and dash-dotted lines indicate  $2m_{\tilde{\chi}_1^0}$ , equal to  $m_{a_1}$ ,  $m_{h_2}$ , and  $m_Z$ , respectively.

between the NLSPs and LSP, the integrated luminosity  $300 \text{ fb}^{-1}$  is not sufficient. Hence, we attempted to increase the luminosity to  $3000 \text{ fb}^{-1}$ , the result is that nearly all samples can be checked with  $ss > 5$  at  $3000 \text{ fb}^{-1}$ .

In Fig. 4, we show the signal significance  $ss$  on the planes of  $m_{\tilde{\chi}_1^0}$  versus  $m_{a_1}$ . We can see that there are mainly two mechanisms for dark matter annihilation: the  $a_1$  funnel, where  $2m_{\tilde{\chi}_1^0} \approx m_{a_1}$ , and the  $h_2/Z$  funnel where  $2m_{\tilde{\chi}_1^0} \approx m_{h_2}$  or  $2m_{\tilde{\chi}_1^0} \approx m_Z$ . Unfortunately, searching for the NLSPs does not contribute to distinguishing these two

mechanisms. Meanwhile, as shown in Ref. [26], the spin-independent cross-section of  $\tilde{\chi}_1^0$  can be sizable, such that singlino-dominated dark matter may be accessible in the future direct detections, such as XENONnT and LUX-ZEPLIN (LZ-7 2T).

## 4 Conclusions

We discussed the light higgsino-dominated NLSPs in the scNMSSM, which is also called the non-universal Higgs mass (NUHM) version of the NMSSM. We use the scenario with singlino-dominated  $\tilde{\chi}_1^0$  and SM-like  $h_2$  in the scan result of our previous study on scNMSSM, where we considered the constraints including theoretical constraints of vacuum stability and the Landau pole, experimental constraints of Higgs data, muon  $g-2$ ,  $B$  physics, dark matter relic density and direct searches, etc. In our scenario, the bino and wino are heavy because of the high mass bound of gluino and the unification of gaugino masses at the GUT scale. Thus, the  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_{2,3}^0$  are higgsino-dominated and mass-degenerated NLSPs.

We first investigate the direct constraints to these light higgsino-dominated NLSPs, including searching for the SUSY particle at the LHC Run-I and Run-II. We use Monte Carlo algorithm to perform detailed simulations to impose these constraints from the search of SUSY particles at the LHC. Then, we discuss the possibility of checking the light higgsino-dominated NLSPs at the HL-

LHC in the future. We use the same analysis by increasing the integrated luminosity to  $300 \text{ fb}^{-1}$  and  $3000 \text{ fb}^{-1}$ .

Finally, we come to the following conclusions regarding the higgsino-dominated  $100 \sim 200 \text{ GeV}$  NLSPs in scNMSSM:

- Among the search results for electroweakinos, the ‘multi-lepton final state’ constrains our scenario the most, and it can exclude some of our samples. Meanwhile, with all data at Run I and up to  $36 \text{ fb}^{-1}$  data at Run II at the LHC, the search results by ATLAS and CMS still cannot exclude the light higgsino-dominated NLSPs of  $100 \sim 200 \text{ GeV}$ .
- When the mass difference with  $\tilde{\chi}_1^0$  is smaller than  $m_{h_2}$ ,  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_3^0$  have different preferences on decaying to  $Z/Z^*$  or  $h_1$ .
- The best channels to detect the NLSPs are though the real two-body decay  $\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 W$  and  $\tilde{\chi}_{2,3}^0 \rightarrow \tilde{\chi}_1^0 Z/h_2$ . When the mass difference is sufficient, most of the samples can be studied at the  $5 \sigma$  level with future  $300 \text{ fb}^{-1}$  data at the LHC. For  $3000 \text{ fb}^{-1}$  data at the LHC, nearly all of the samples can be studied at the  $5\sigma$  level even if the mass difference is insufficient.
- The  $a_1$  funnel and the  $h_2/Z$  funnel are the two main mechanisms for the singlino-dominated LSP annihilating, which cannot be distinguished by searching for NLSPs.

*We thank Yuanfang Yue, Yang Zhang, and Liangliang Shang for useful discussions.*

## References

- J. E. Kim and H. P. Nilles, Phys. Lett., **138B**: 150 (1984)
- J. Cao, Z. Heng, J. M. Yang *et al.*, JHEP, **1210**: 079 (2012), arXiv:1207.3698[hep-ph]
- J. J. Cao, Z. X. Heng, J. M. Yang *et al.*, JHEP, **1203**: 086 (2012), arXiv:1202.5821[hep-ph]
- J. Cao, Z. Heng, D. Li *et al.*, Phys. Lett. B, **710**: 665 (2012), arXiv:1112.4391[hep-ph]
- J. Ellis and K. A. Olive, Eur. Phys. J. C, **72**: 2005 (2012), arXiv:1202.3262[hep-ph]
- F. Wang, K. Wang, J. M. Yang *et al.*, JHEP, **1812**: 041 (2018), arXiv:1808.10851[hep-ph]
- C. Han, K. i. Hikasa, L. Wu *et al.*, Phys. Lett. B, **769**: 470 (2017), arXiv:1612.02296[hep-ph]
- P. Bechtle *et al.*, Eur. Phys. J. C, **76**(2): 96 (2016), arXiv:1508.05951[hep-ph]
- P. Athron *et al.* (GAMBIT Collaboration), Eur. Phys. J. C, **77**(12): 824 (2017), arXiv:1705.07935[hep-ph]
- K. Kowalska, S. Munir, L. Roszkowski *et al.*, Phys. Rev. D, **87**: 115010 (2013), arXiv:1211.1693[hep-ph]
- J. F. Gunion, Y. Jiang, and S. Kraml, Phys. Lett. B, **710**: 454 (2012), arXiv:1201.0982[hep-ph]
- U. Ellwanger and C. Hugonie, Adv. High Energy Phys., **2012**: 625389 (2012), arXiv:1203.5048[hep-ph]
- C. Beskidt, W. de Boer, and D. I. Kazakov, Phys. Lett. B, **726**: 758 (2013), arXiv:1308.1333[hep-ph]
- D. Kim, P. Athron, C. Balazs *et al.*, Phys. Rev. D, **90**(5): 055008 (2014), arXiv:1312.4150[hep-ph]
- A. Fowlie, Eur. Phys. J. C, **74**(10): 3105 (2014), arXiv:1407.7534[hep-ph]
- D. G. Cerdeno, V. De Romeri, V. Martin-Lozano *et al.*, Eur. Phys. J. C, **78**(4): 290 (2018), arXiv:1707.03990[hep-ph]
- U. Ellwanger, A. Florent, and D. Zerwas, JHEP, **1101**: 103 (2011), arXiv:1011.0931[hep-ph]
- G. Panotopoulos, J. Phys. Conf. Ser., **259**: 012064 (2010), arXiv:1010.4481[hep-ph]
- D. E. Lopez-Fogliani, L. Roszkowski, R. Ruiz de Austri *et al.*, Phys. Rev. D, **80**: 095013 (2009), arXiv:0906.4911[hep-ph]
- G. Belanger, C. Hugonie, and A. Pukhov, JCAP, **0901**: 023 (2009), arXiv:0811.3224[hep-ph]
- A. Djouadi, U. Ellwanger, and A. M. Teixeira, JHEP, **0904**: 031 (2009), arXiv:0811.2699[hep-ph]
- U. Ellwanger, AIP Conf. Proc., **1078**: 73 (2009), arXiv:0809.0779[hep-ph]
- C. Hugonie, G. Belanger, and A. Pukhov, JCAP, **0711**: 009 (2007), arXiv:0707.0628[hep-ph]
- D. Das, U. Ellwanger, and A. M. Teixeira, JHEP, **1304**: 117 (2013), arXiv:1301.7584[hep-ph]
- U. Ellwanger and C. Hugonie, JHEP, **1408**: 046 (2014), arXiv:1405.6647[hep-ph]
- K. Wang, F. Wang, J. Zhu *et al.*, Chin. Phys. C, **42**(10): 103109 (2018), arXiv:1811.04435[hep-ph]
- W. Abdallah, A. Chatterjee, and A. Datta, JHEP, **1909**: 095 (2019), arXiv:1907.06270[hep-ph]
- D. Das, U. Ellwanger, and A. M. Teixeira, JHEP, **1204**: 067 (2012), arXiv:1203.5048[hep-ph]

- (2012), arXiv: 1202.5244[hep-ph]
- 29 A. Pyarelal and S. Su, arXiv: 1907.11326[hep-ph]
  - 30 U. Ellwanger and C. Hugonie, *Eur. Phys. J. C*, **78**(9): 735 (2018), arXiv: 1806.09478[hep-ph]
  - 31 U. Ellwanger, *JHEP*, **1702**: 051 (2017), arXiv: 1612.06574[hep-ph]
  - 32 Q. F. Xiang, X. J. Bi, P. F. Yin *et al.*, *Phys. Rev. D*, **94**(5): 055031 (2016), arXiv: 1606.02149[hep-ph]
  - 33 M. Badziak, M. Olechowski, and P. Szczerbiak, *JHEP*, **1603**: 179 (2016), arXiv: 1512.02472[hep-ph]
  - 34 U. Ellwanger and A. M. Teixeira, *JHEP*, **1410**: 113 (2014), arXiv: 1406.7221[hep-ph]
  - 35 U. Ellwanger, *JHEP*, **1311**: 108 (2013), arXiv: 1309.1665[hep-ph]
  - 36 F. Domingo, J. S. Kim, V. Martin-Lozano *et al.*, arXiv: 1812.05186[hep-ph]
  - 37 A. Titterton, U. Ellwanger, H. U. Flaecher *et al.*, *JHEP*, **1810**: 064 (2018), arXiv: 1807.10672[hep-ph]
  - 38 J. S. Kim and T. S. Ray, *Eur. Phys. J. C*, **75**: 40 (2015), arXiv: 1405.3700[hep-ph]
  - 39 B. Dutta, Y. Gao, and B. Shakya, *Phys. Rev. D*, **91**(3): 035016 (2015), arXiv: 1412.2774[hep-ph]
  - 40 J. Cao, L. Shang, P. Wu *et al.*, *Phys. Rev. D*, **91**(5): 055005 (2015), arXiv: 1410.3239[hep-ph]
  - 41 R. Enberg, S. Munir, C. Perez de los Heros *et al.*, arXiv: 1506.05714[hep-ph]
  - 42 C. T. Potter, *Eur. Phys. J. C*, **76**(1): 44 (2016), arXiv: 1505.05554[hep-ph]
  - 43 J. Cao, Y. He, L. Shang *et al.*, *JHEP*, **1608**: 037 (2016), arXiv: 1606.04416[hep-ph]
  - 44 J. Cao, Y. He, L. Shang *et al.*, *Phys. Rev. D*, **99**(7): 075020 (2019), arXiv: 1810.09143[hep-ph]
  - 45 U. Ellwanger, C. Hugonie, and A. M. Teixeira, *Phys. Rept.*, **496**: 1 (2010), arXiv: 0910.1785[hep-ph]
  - 46 D. J. Miller, R. Nevzorov, and P. M. Zerwas, *Nucl. Phys. B*, **681**: 3 (2004), arXiv: hep-ph/0304049
  - 47 J. Cao, L. Shang, P. Wu *et al.*, *JHEP*, **1510**: 030 (2015), arXiv: 1506.06471[hep-ph]
  - 48 M. Badziak, M. Olechowski, and P. Szczerbiak, *JHEP*, **1603**: 179 (2016), arXiv: 1512.02472[hep-ph]
  - 49 D. Das, U. Ellwanger, and A. M. Teixeira, *JHEP*, **1204**: 067 (2012), arXiv: 1202.5244[hep-ph]
  - 50 W. Abdallah, A. Chatterjee, and A. Datta, *JHEP*, **1909**: 095 (2019), arXiv: 1907.06270[hep-ph]
  - 51 M. Drees, H. Dreiner, D. Schmeier *et al.*, *Comput. Phys. Commun.*, **187**: 227 (2015), arXiv: 1312.2591[hep-ph]
  - 52 J. S. Kim, D. Schmeier, J. Tattersall *et al.*, *Comput. Phys. Commun.*, **196**: 535 (2015), arXiv: 1503.01123[hep-ph]
  - 53 D. Dereks, N. Desai, J. S. Kim *et al.*, *Comput. Phys. Commun.*, **221**: 383 (2017), arXiv: 1611.09856[hep-ph]
  - 54 U. Ellwanger and C. Hugonie, *Comput. Phys. Commun.*, **177**: 399 (2007), [hep-ph/0612134]
  - 55 D. Das, U. Ellwanger, and A. M. Teixeira, *Comput. Phys. Commun.*, **183**: 774 (2012), arXiv: 1106.5633[hep-ph]
  - 56 U. Ellwanger, J. F. Gunion, and C. Hugonie, *JHEP*, **0502**: 066 (2005), [hep-ph/0406215]
  - 57 U. Ellwanger and C. Hugonie, *Comput. Phys. Commun.*, **175**: 290 (2006), [hep-ph/0508022]
  - 58 P. Bechtle, O. Brein, S. Heinemeyer *et al.*, *Eur. Phys. J. C*, **74**(3): 2693 (2014), arXiv: 1311.0055[hep-ph]
  - 59 F. Ambrogio *et al.*, *Comput. Phys. Commun.*, **227**: 72 (2018), arXiv: 1701.06586[hep-ph]
  - 60 S. Kraml, S. Kulkarni, U. Laa *et al.*, *Eur. Phys. J. C*, **74**: 2868 (2014), arXiv: 1312.4175[hep-ph]
  - 61 (CMS Collaboration), *JHEP*, **1803**: 166 (2018), arXiv: 1709.05406[hep-ex]
  - 62 (CMS Collaboration), *JHEP*, **1803**: 160 (2018), arXiv: 1801.03957[hep-ex]
  - 63 (ATLAS Collaboration), ATLASCONF-2016-096.
  - 64 (CMS Collaboration), CMS-PAS-SUS-16-025.
  - 65 (CMS Collaboration), *Phys. Lett. B*, **782**: 440 (2018), arXiv: 1801.01846[hep-ex]
  - 66 (ATLAS Collaboration), ATLAS-CONF-2016-013, ATLASCONF-2016-050, ATLAS-CONF-2016-054, ATLAS-CONF-2016-076, ATLAS-CONF-2015-082, arXiv: 1712.08119, arXiv: 1712.02332, arXiv: 1709.04183, arXiv: 1708.07875, arXiv: 1706.03731
  - 67 J. Alwall *et al.*, *JHEP*, **1407**: 079 (2014), arXiv: 1405.0301[hep-ph]
  - 68 W. Beenakker, R. Hopker, and M. Spira, hep-ph/9611232.
  - 69 T. Sjöstrand *et al.*, *Comput. Phys. Commun.*, **191**: 159 (2015), arXiv: 1410.3012[hep-ph]
  - 70 (DELPHES 3 Collaboration), *JHEP*, **1402**: 057 (2014), arXiv: 1307.6346[hep-ex]
  - 71 M. Cacciari, G. P. Salam, and G. Soyez, *JHEP*, **0804**: 063 (2008), arXiv: 0802.1189[hep-ph]
  - 72 (ATLAS Collaboration), arXiv: 1912.08479 [hep-ex]; (ATLAS Collaboration), ATLAS-CONF-2019-020.
  - 73 C. G. Lester and D. J. Summers, *Phys. Lett. B*, **463**: 99 (1999), [hep-ph/9906349]
  - 74 A. Barr, C. Lester, and P. Stephens, *J. Phys. G*, **29**: 2343 (2003), [hep-ph/0304226]; H. C. Cheng and Z. Han, *JHEP*, **0812**: 063 (2008), arXiv: 0810.5178[hep-ph]
  - 75 Y. Bai, H. C. Cheng, J. Gallicchio *et al.*, *JHEP*, **1207**: 110 (2012), arXiv: 1203.4813[hep-ph]
  - 76 (ATLAS Collaboration), arXiv: 1908.08215[hep-ex]
  - 77 (ATLAS Collaboration), ATLAS-CONF-2019-008
  - 78 (ATLAS Collaboration), ATLAS-CONF-2019-014
  - 79 (CMS Collaboration), *Phys. Lett. B*, **782**: 440 (2018), arXiv: 1801.01846[hep-ex]
  - 80 (ATLAS Collaboration), ATLAS-CONF-2019-018
  - 81 (CMS Collaboration), CMS-PAS-SUS-18-006
  - 82 S. Buddenbrock *et al.*, *JHEP*, **1910**: 157 (2019), arXiv: 1901.05300[hep-ph]
  - 83 S. von Buddenbrock *et al.*, *Eur. Phys. J. C*, **76**(10): 580 (2016), arXiv: 1606.01674[hep-ph]
  - 84 S. von Buddenbrock *et al.*, *J. Phys. G*, **45**(11): 115003 (2018), arXiv: 1711.07874[hep-ph]
  - 85 (ATLAS Collaboration), ATLAS-CONF-2019-019
  - 86 (CMS Collaboration), CMS-PAS-SUS-18-007
  - 87 (ATLAS Collaboration), *Phys. Rev. D*, **100**(1): 012006 (2019), arXiv: 1812.09432[hep-ex]
  - 88 (ATLAS Collaboration), arXiv: 1909.09226[hep-ex]
  - 89 (ATLAS Collaboration), arXiv: 1911.12606[hep-ex]
  - 90 (CMS Collaboration), arXiv: 1910.01185[hep-ex]